

Polarimetry of planetary atmospheres

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Universiteit Leiden





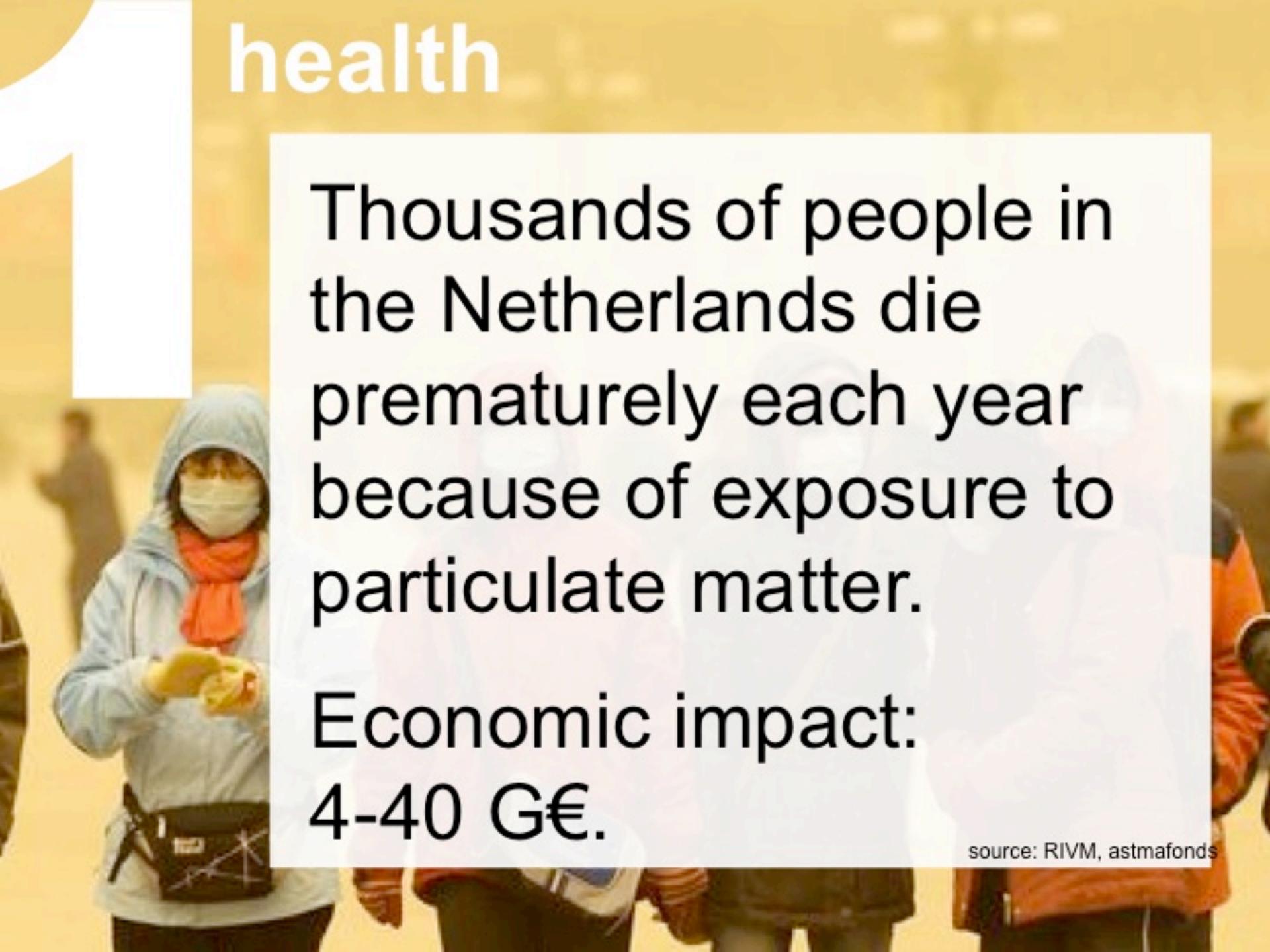
health



health



health



Thousands of people in the Netherlands die prematurely each year because of exposure to particulate matter.

Economic impact:
4-40 G€.

source: RIVM, astmafonds

2

economy



2

economy

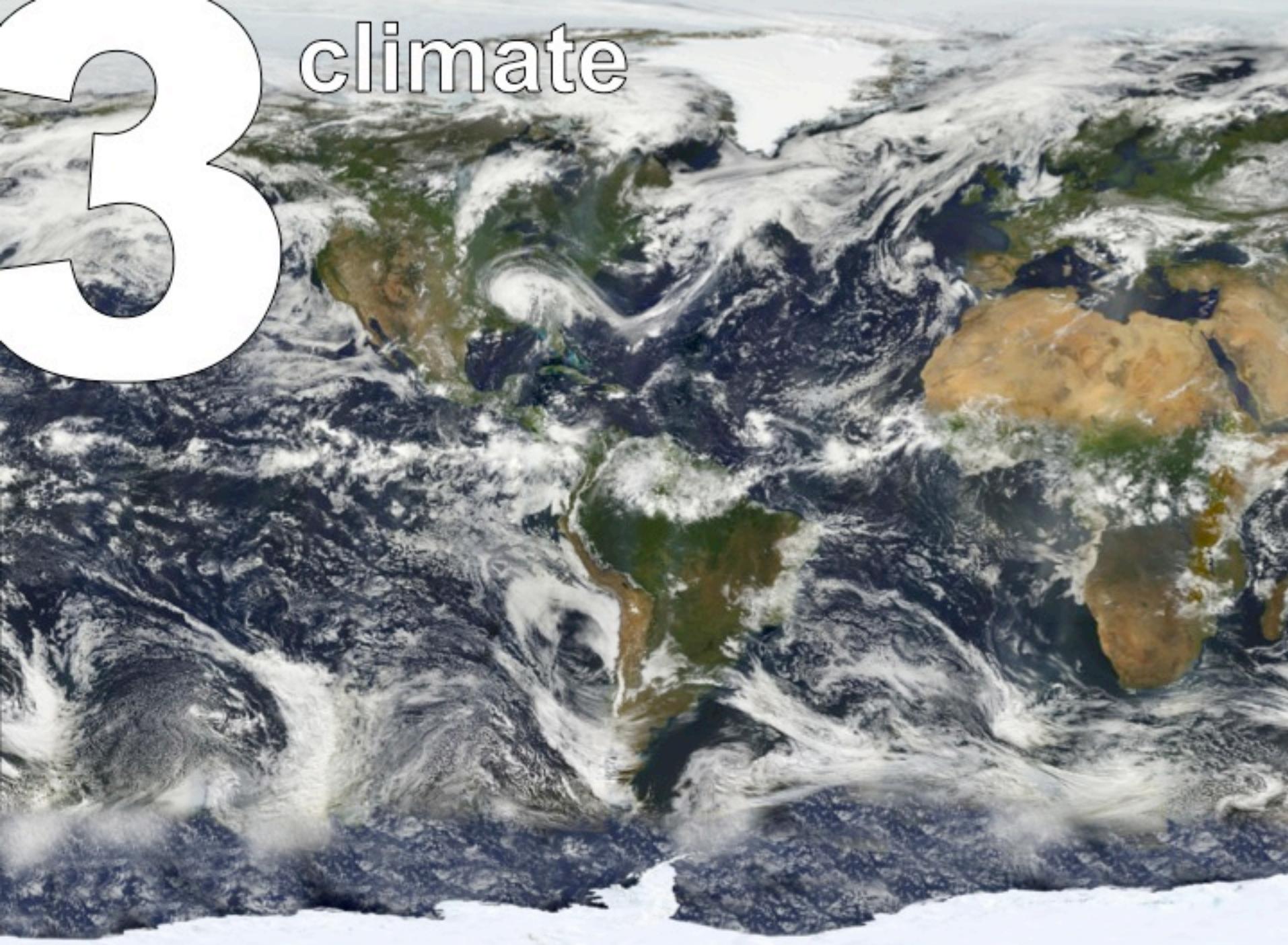




economy

Cancellation of flights due to the volcanic ash clouds in 2010 has cost airlines more than a billion euros.

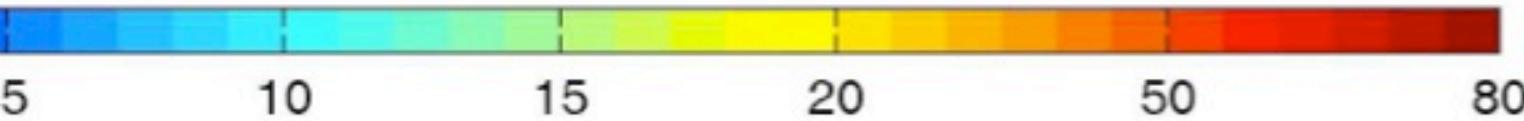
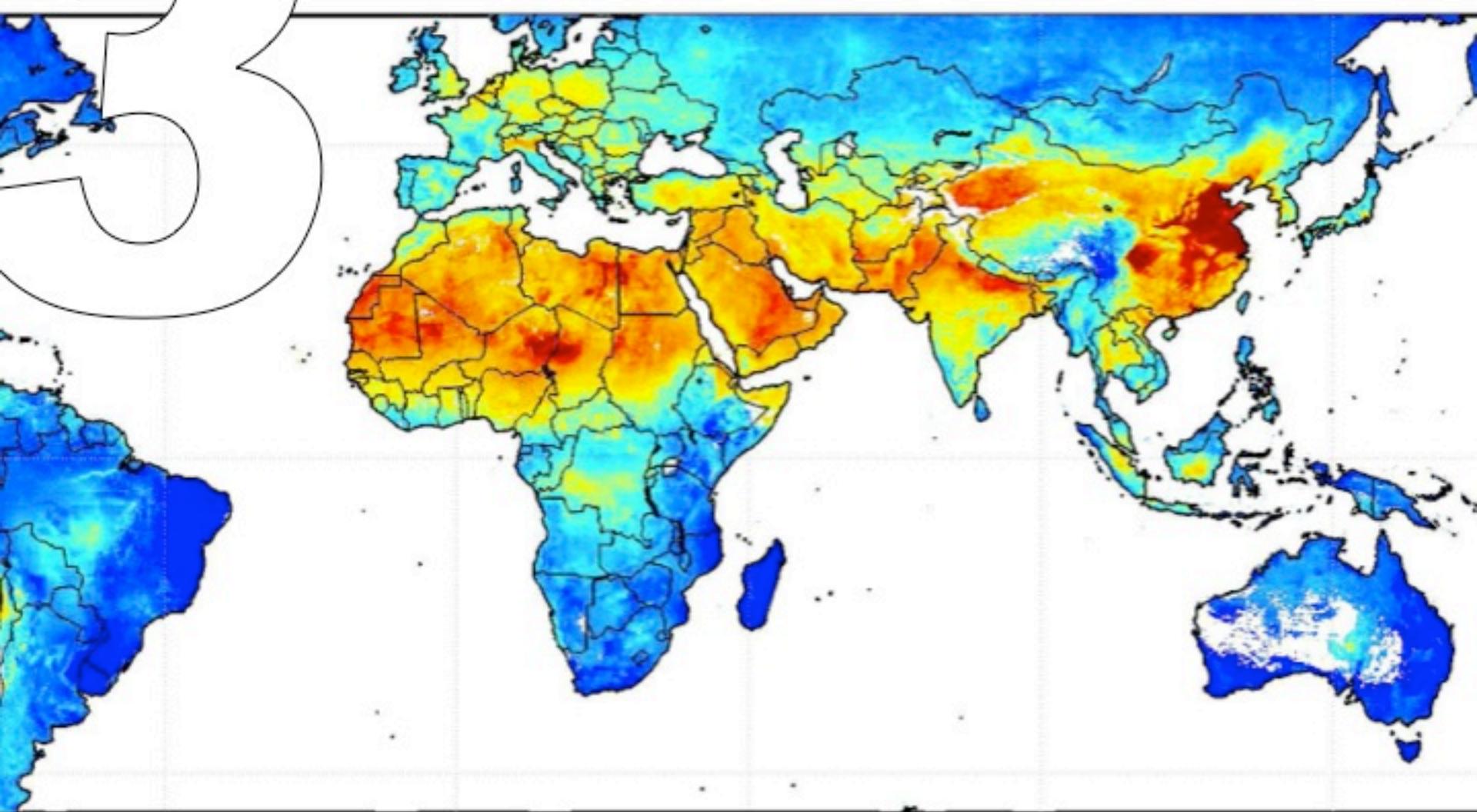
source: IATA



3

climate

climate

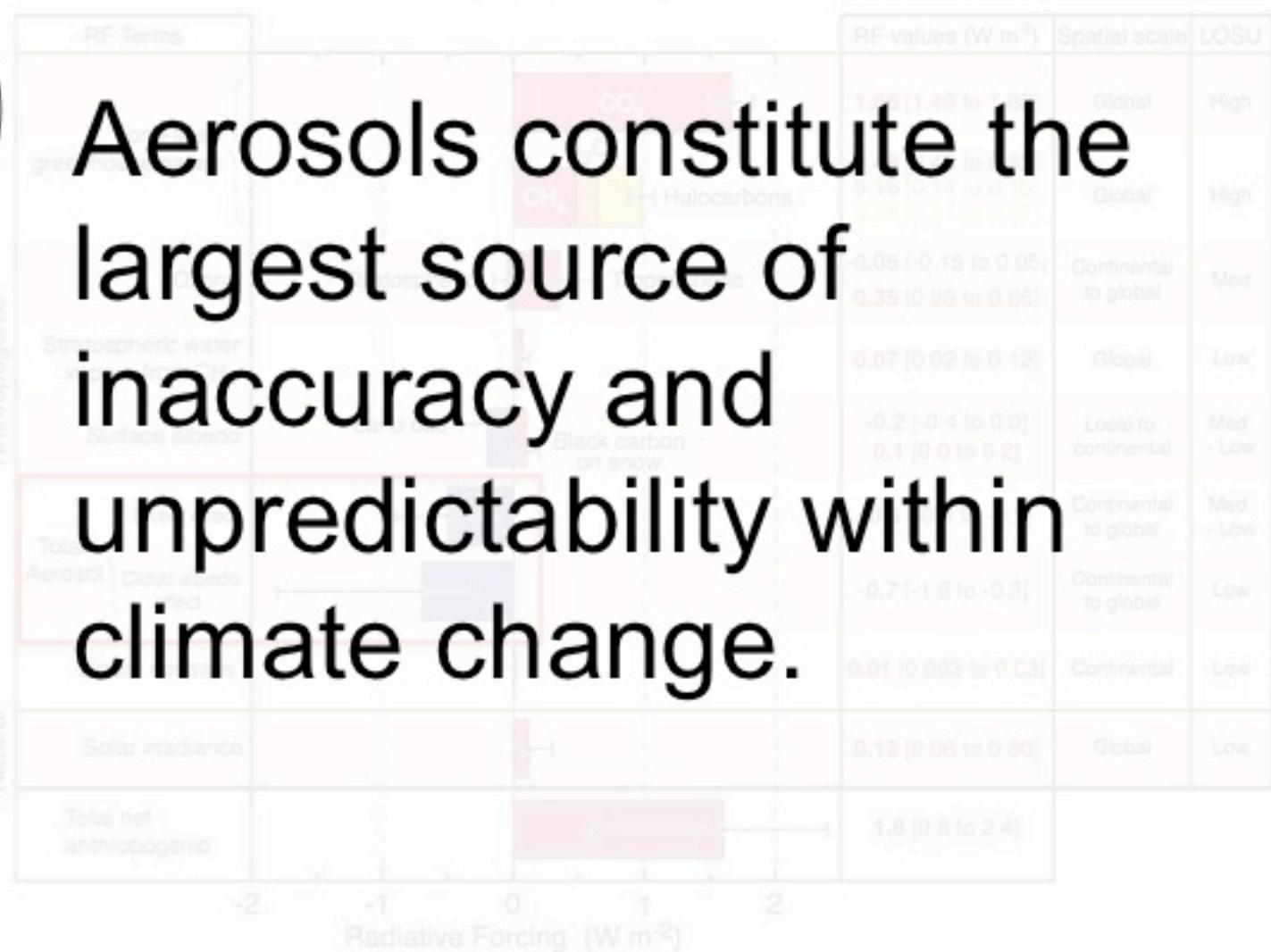


3

climate

Radiative Forcing Components

Aerosols constitute the largest source of inaccuracy and unpredictability within climate change.





spectrum



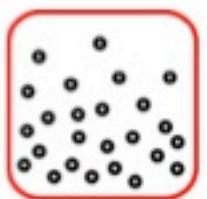
linear polarization



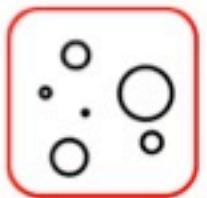
multiple scattering angles



number density



size distribution

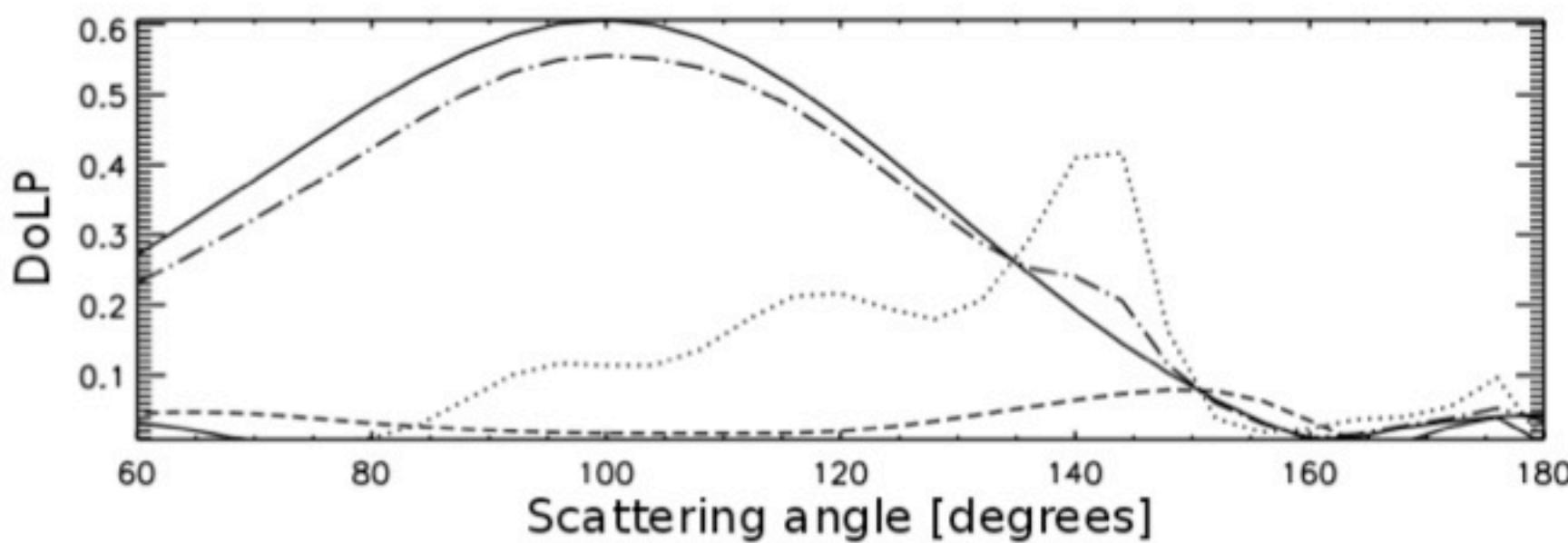
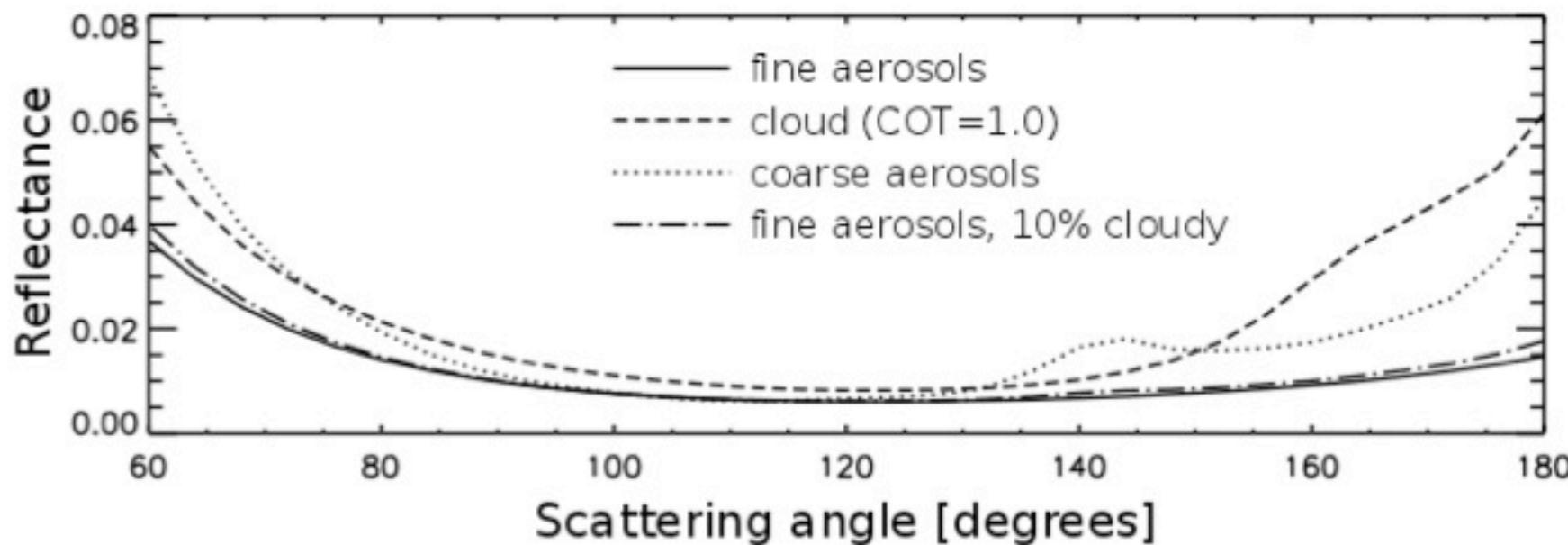


chemical composition

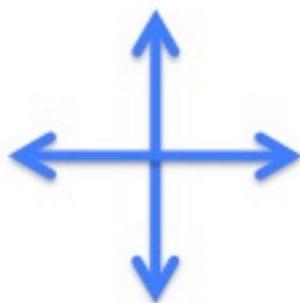


shape





1. temporal modulation



Advantages

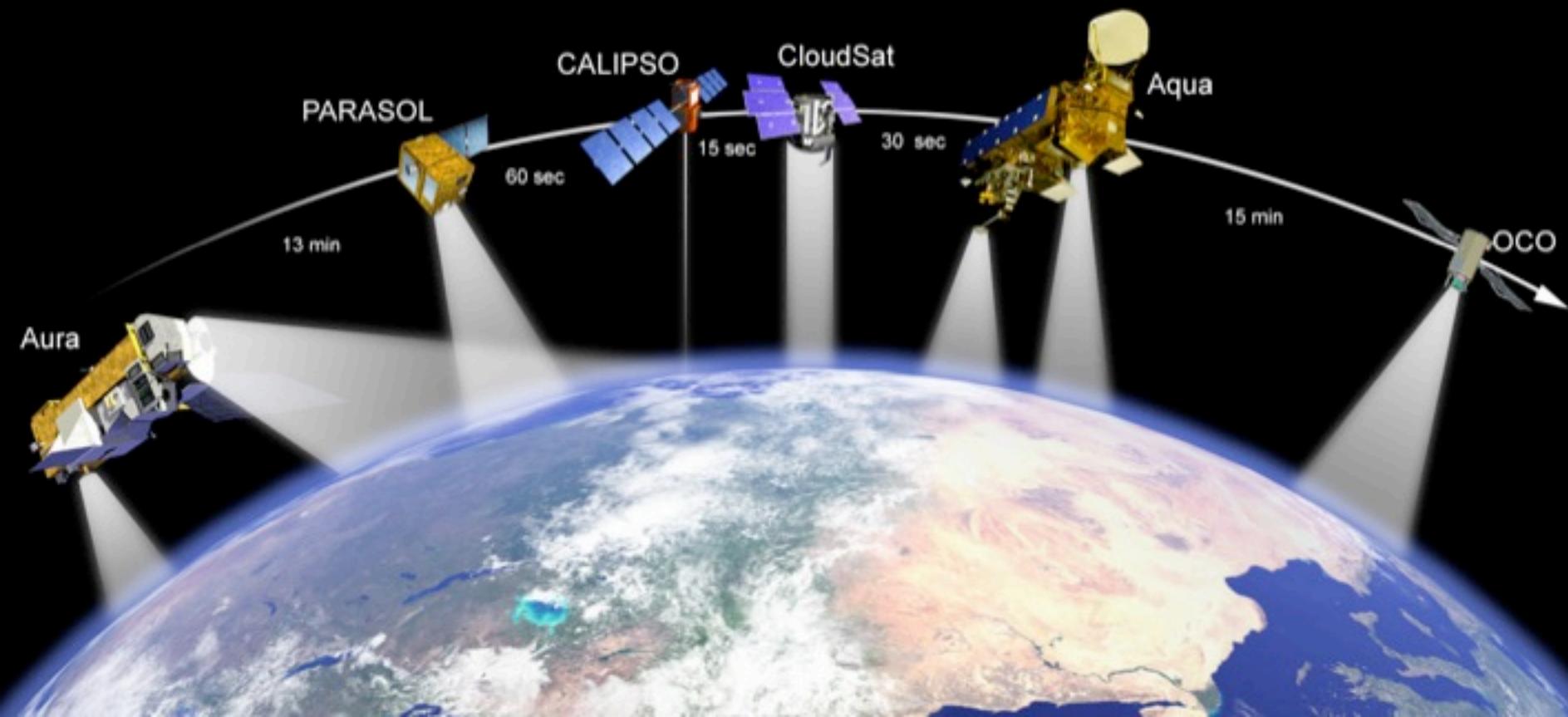
- Single detector
- Easy addition of polarization optics

Disadvantages

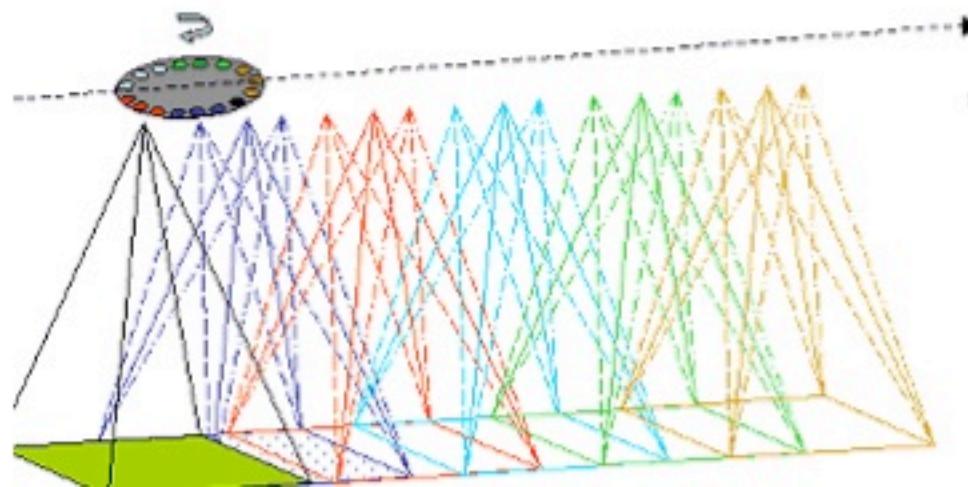
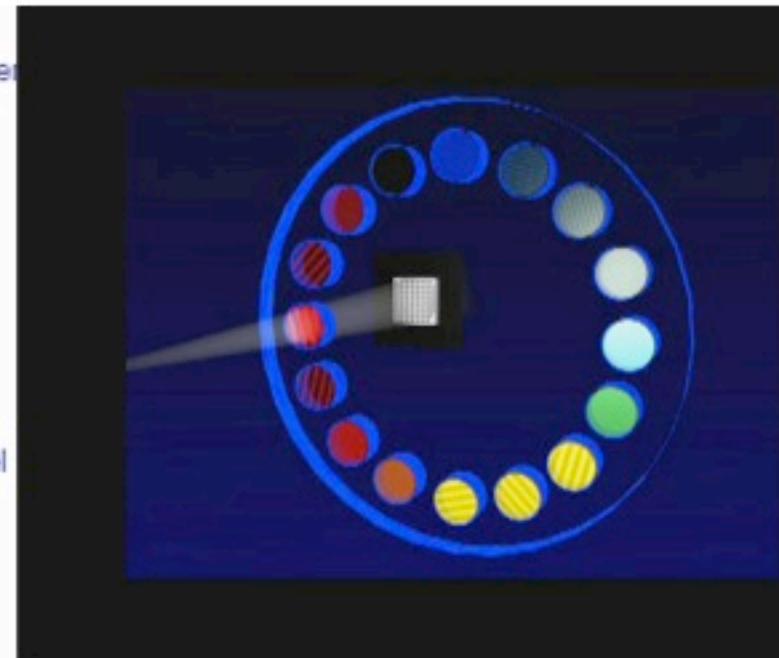
- susceptible to variations in time:
 - flying over scene
 - changing scene
 - atmospheric fluctuations

POLDER/PARASOL

The A-Train

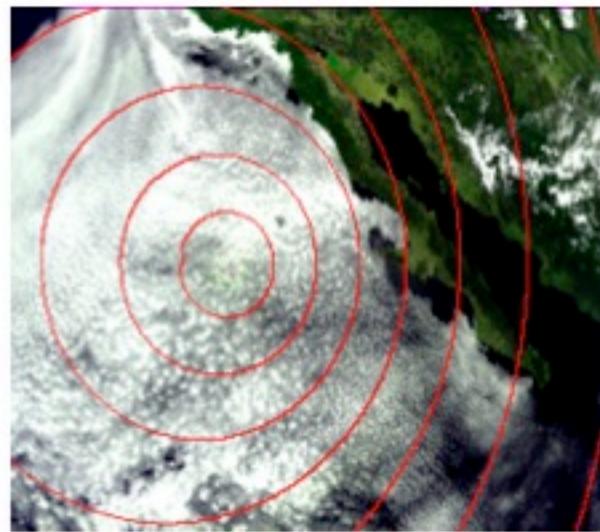


POLDER/PARASOL

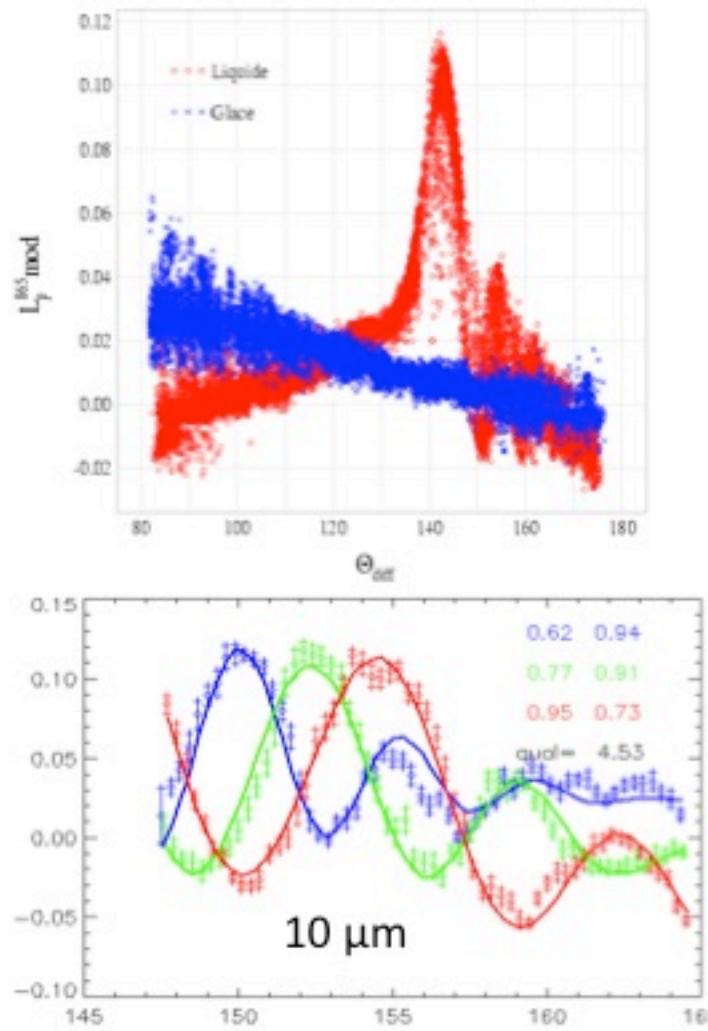
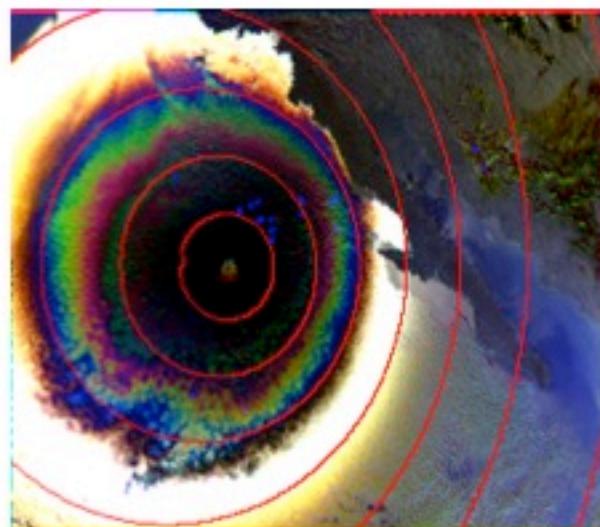


POLDER/PARASOL

intensity
reflectance



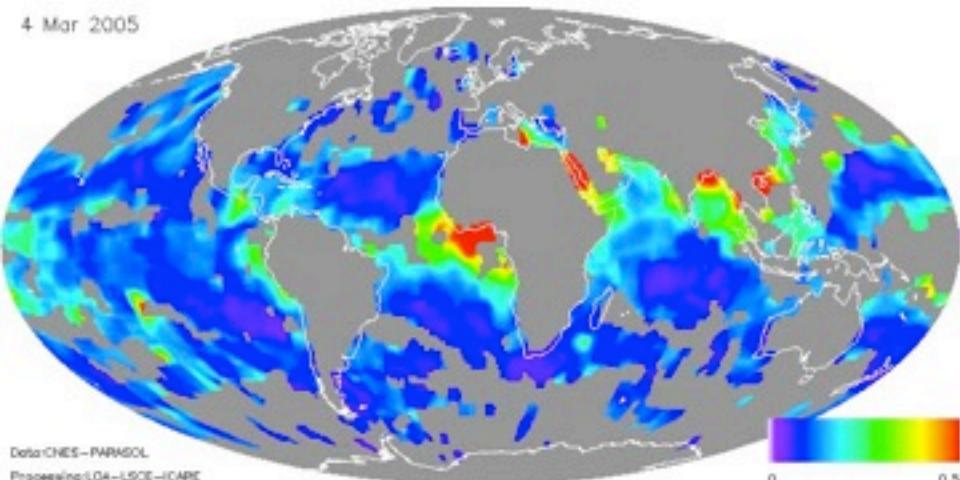
polarized
reflectance



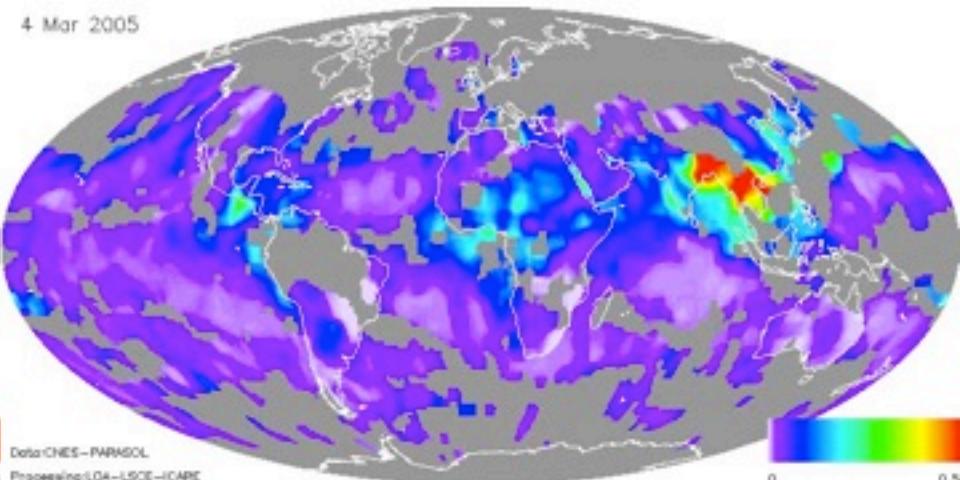
http://smsc.cnes.fr/POLDER/lien3_scie.htm

POLDER/PARASOL

total Aerosol Optical Thickness

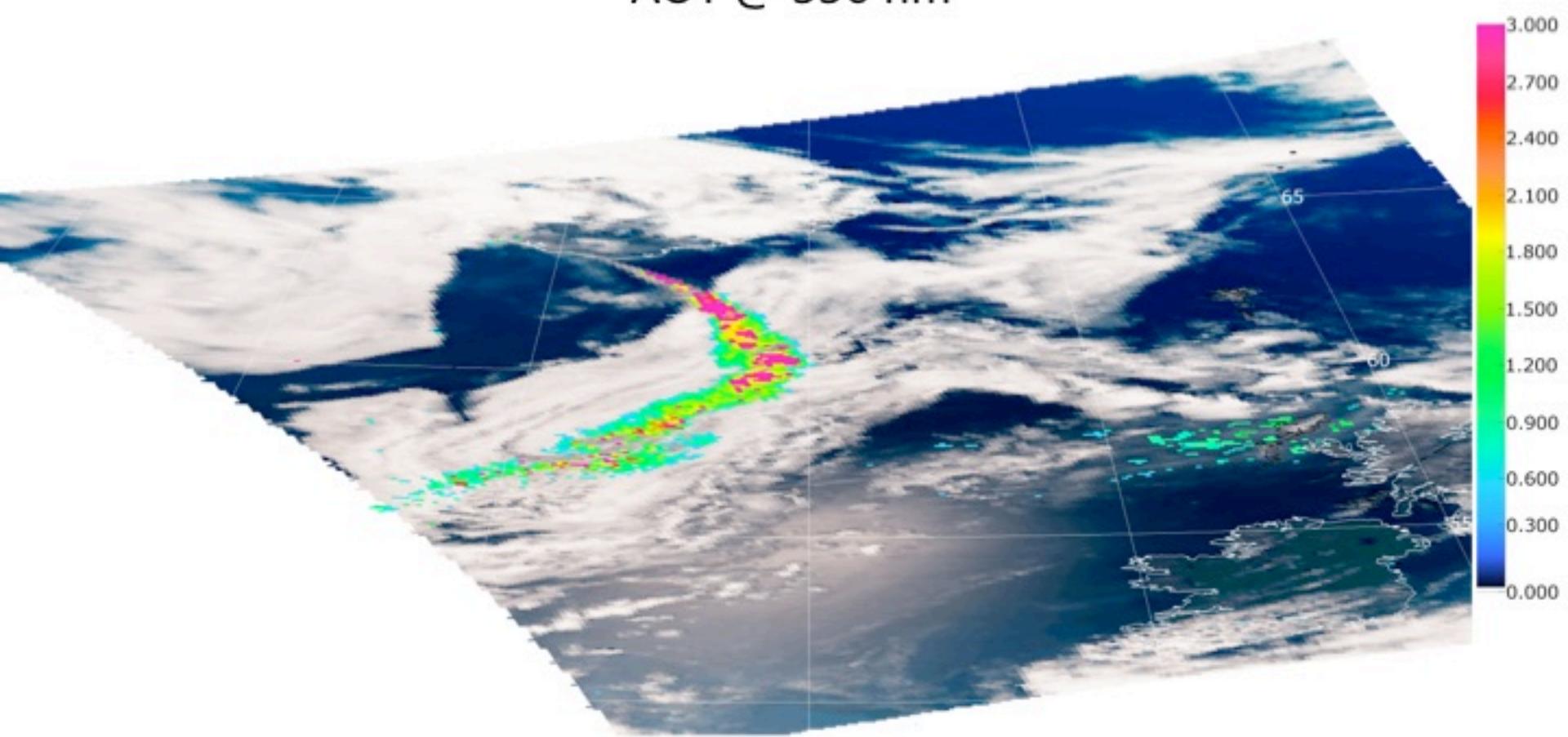


fine mode Aerosol Optical Thickness

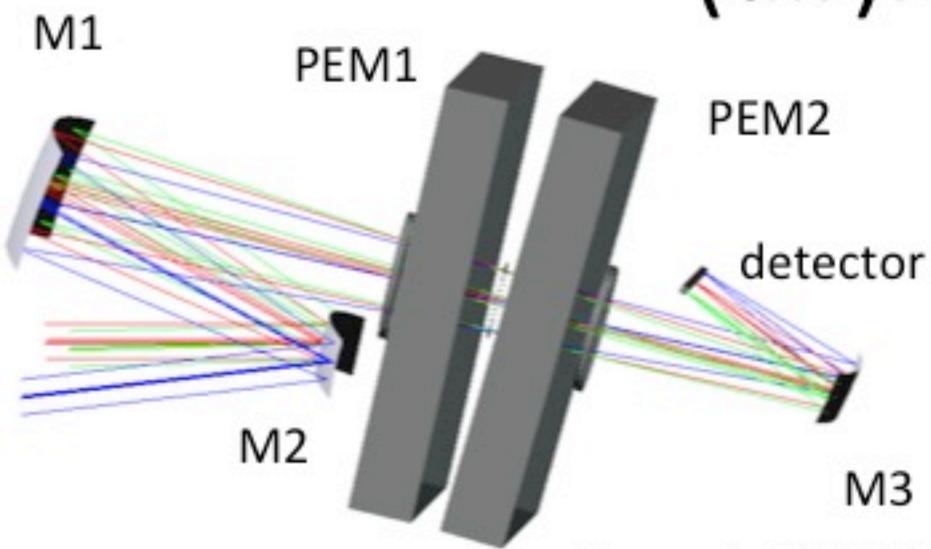


POLDER/PARASOL

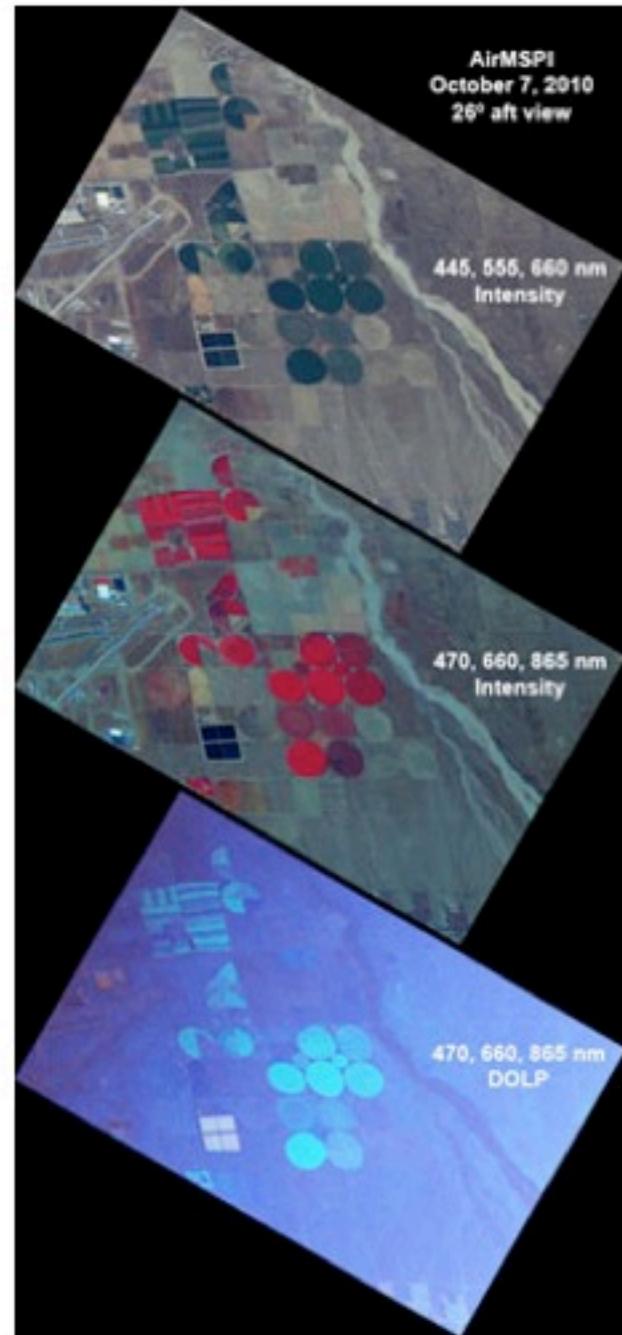
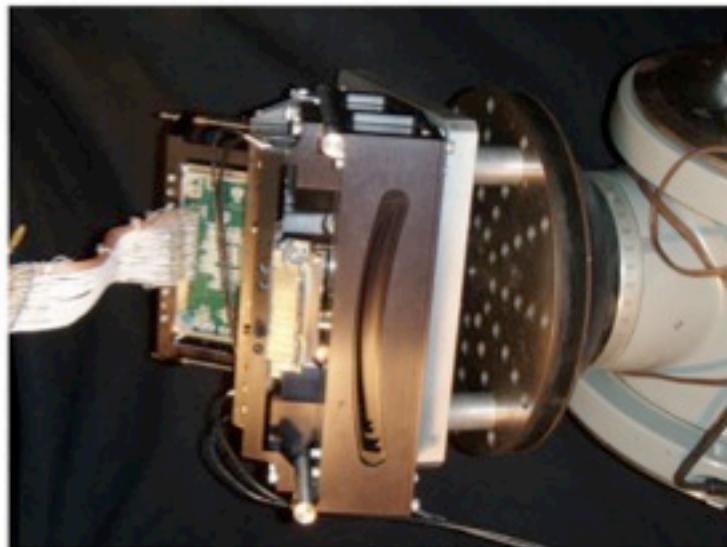
Eyjafjallajokull volcano on May 8, 2010
AOT @ 550 nm



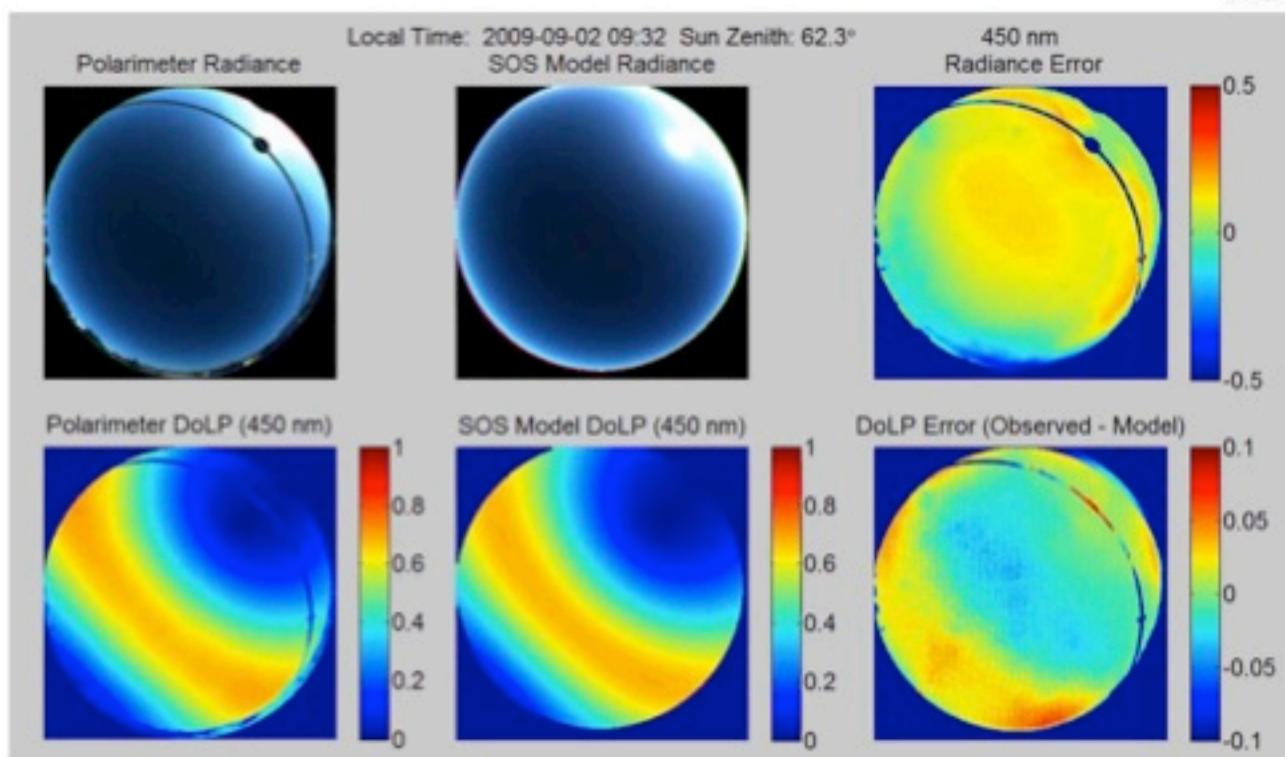
(air)MSPI



Diner et al. (2007, 2010)

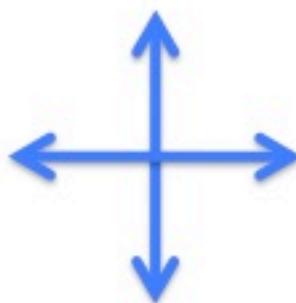


ground-based all sky polarimetry



Pust et al. (2011)

2. spatial modulation



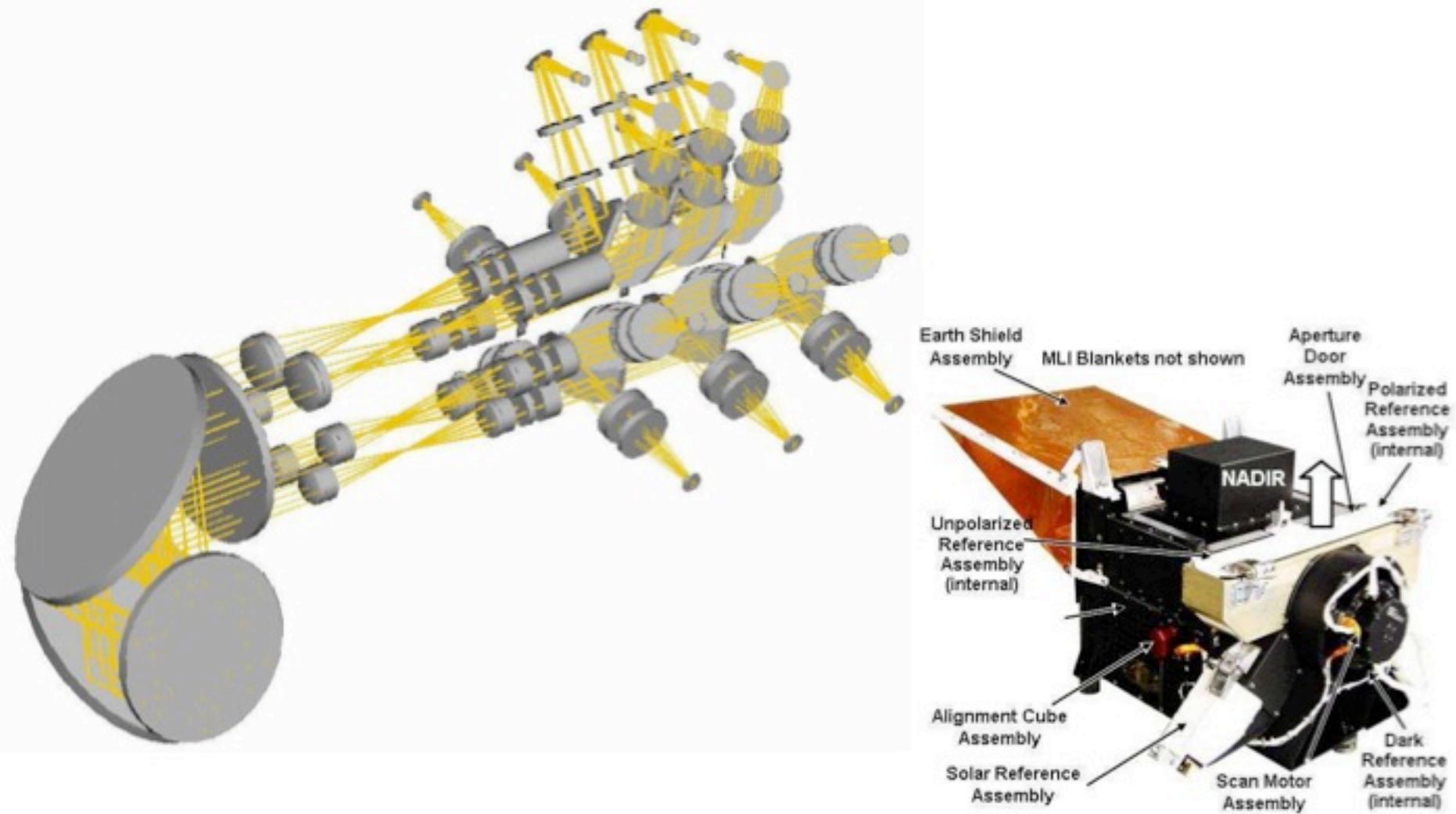
Advantages

- instantaneous measurement

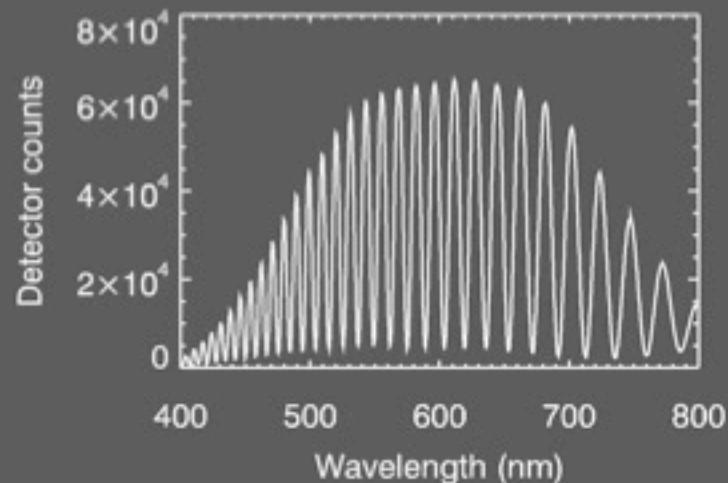
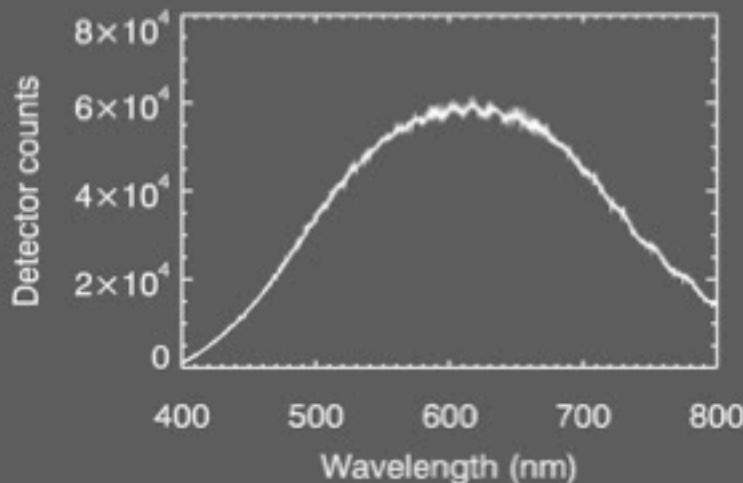
Disadvantages

- susceptible to differences between beams:
 - transmission
 - alignment
 - aberrations
 - detector flat-fielding

GLORY-APS



3. spectral modulation



Advantages

- all info in one shot
- no moving parts
- can be first element

Disadvantages

- spectrograph
- additional pixels



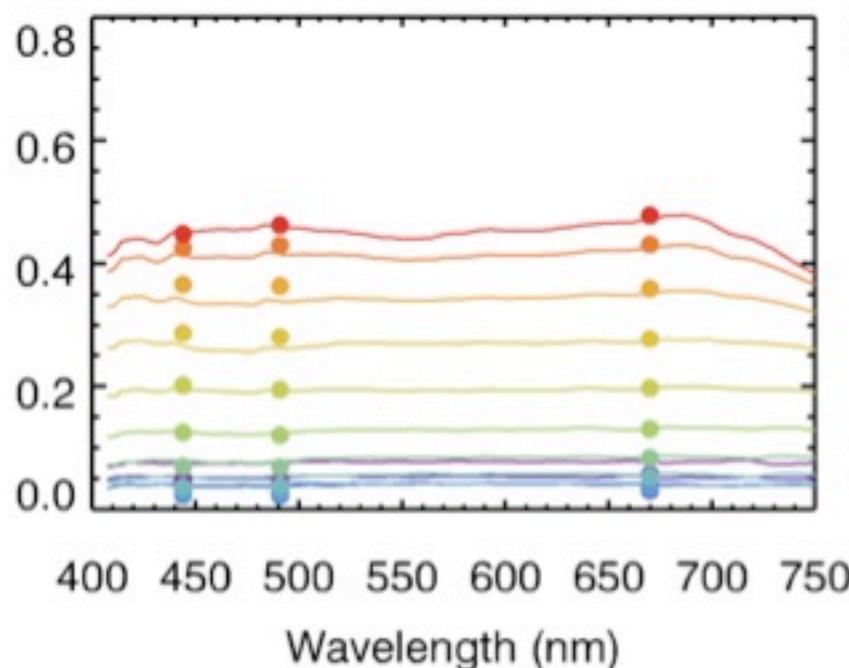


SPEX measurements @ CESAR, May 2012

24-05-2012

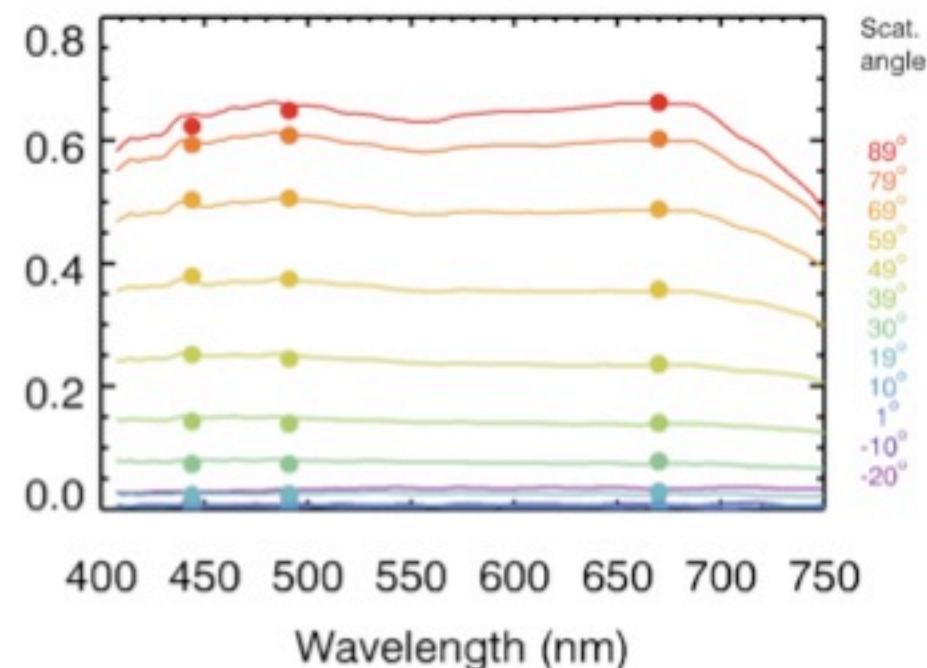
13:08 UTC

DoLP



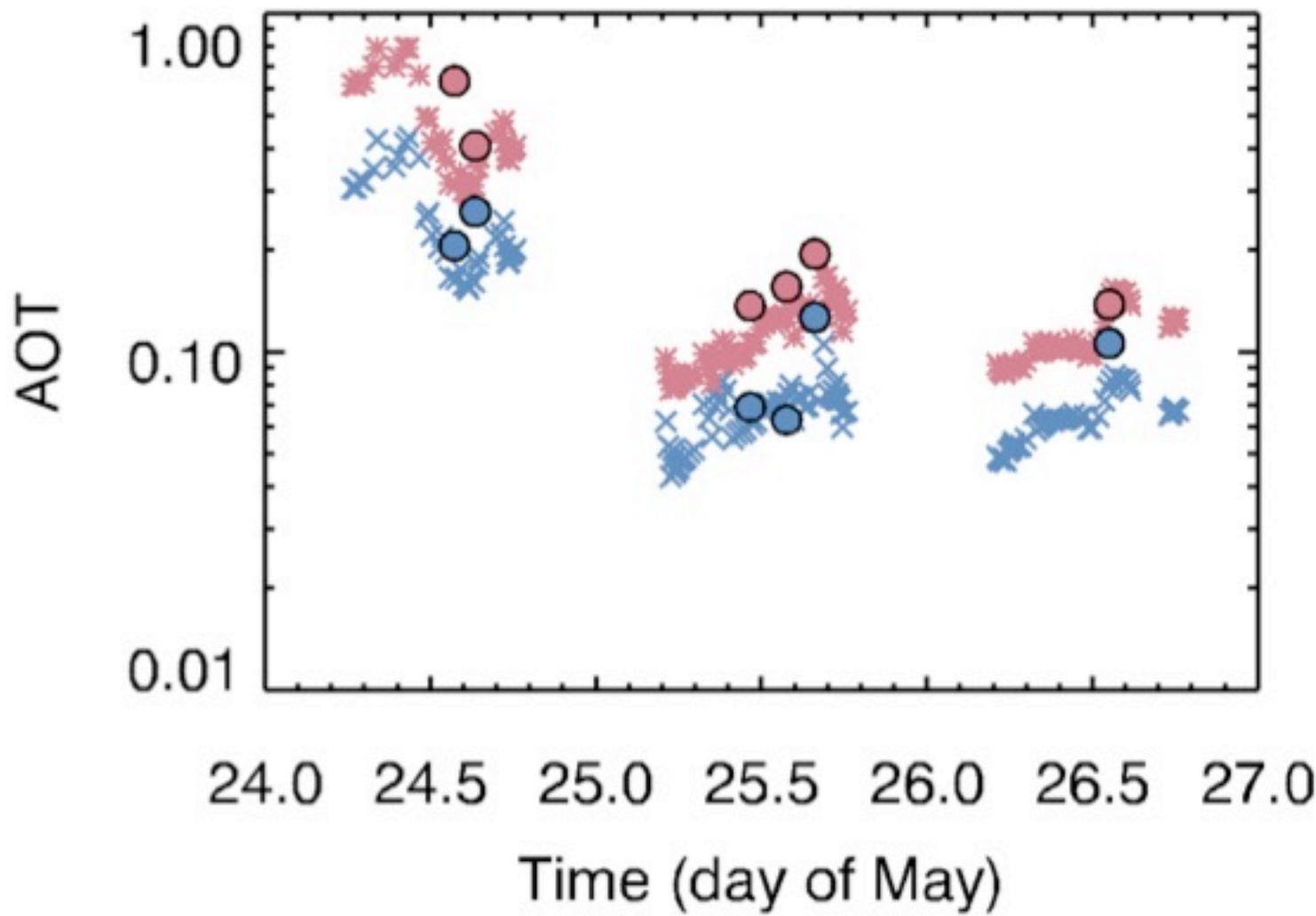
25-05-2012

11:37 UTC

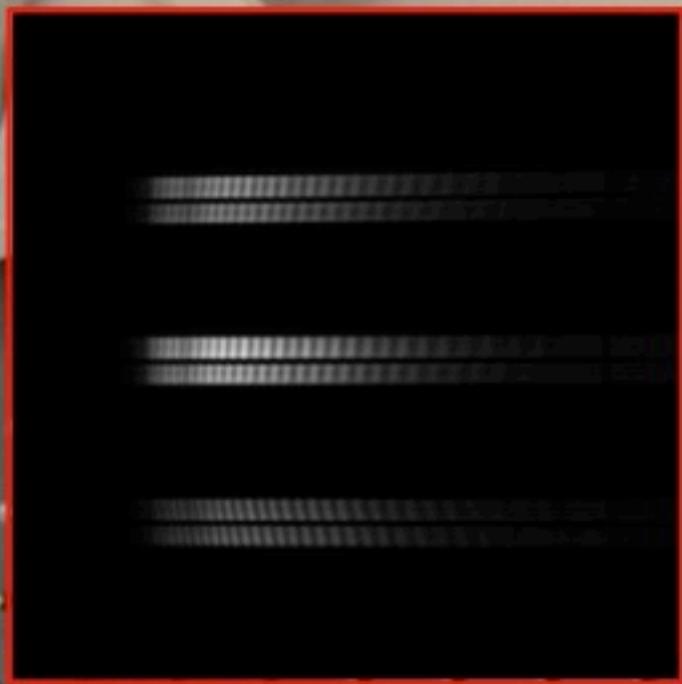
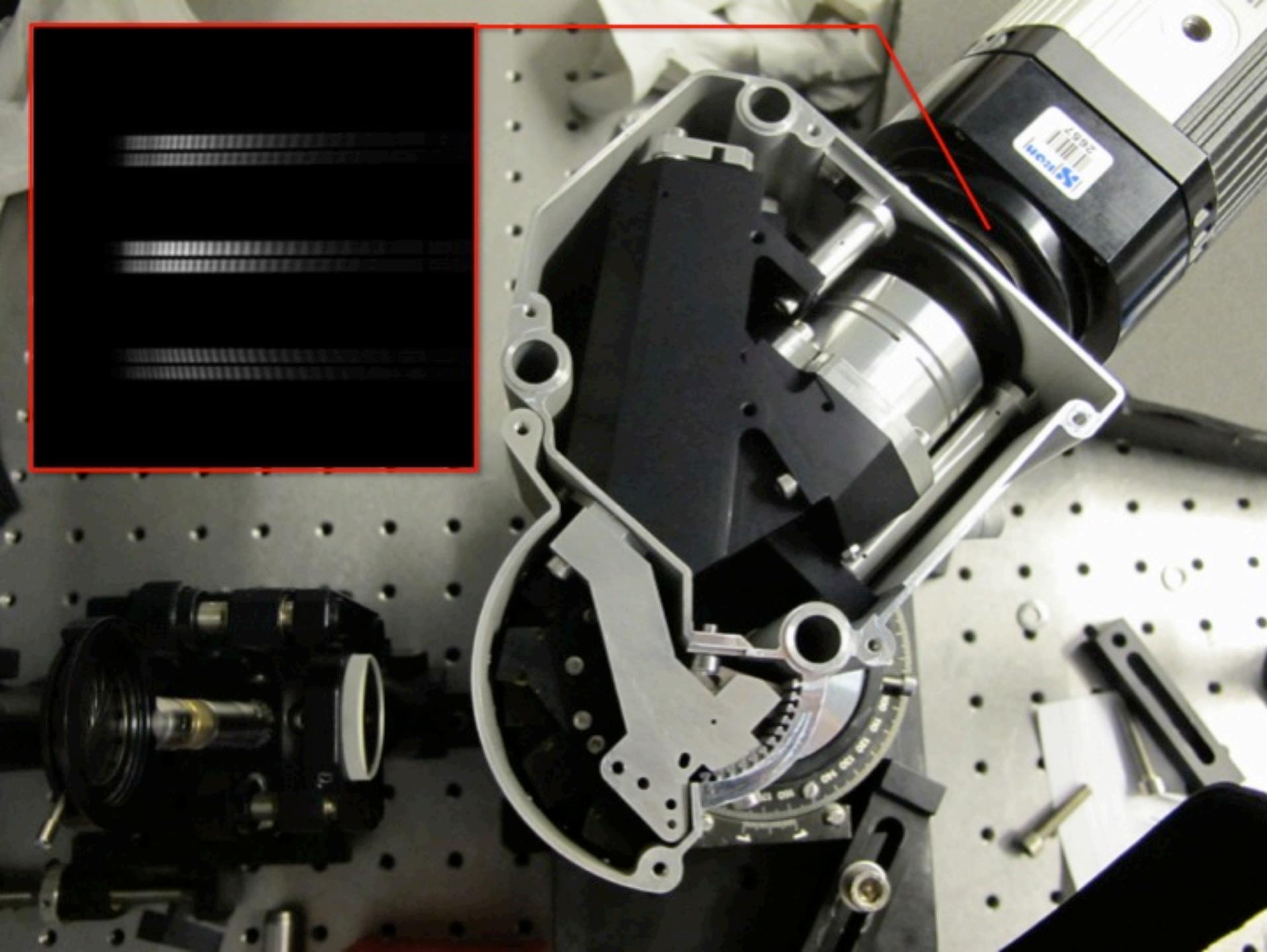


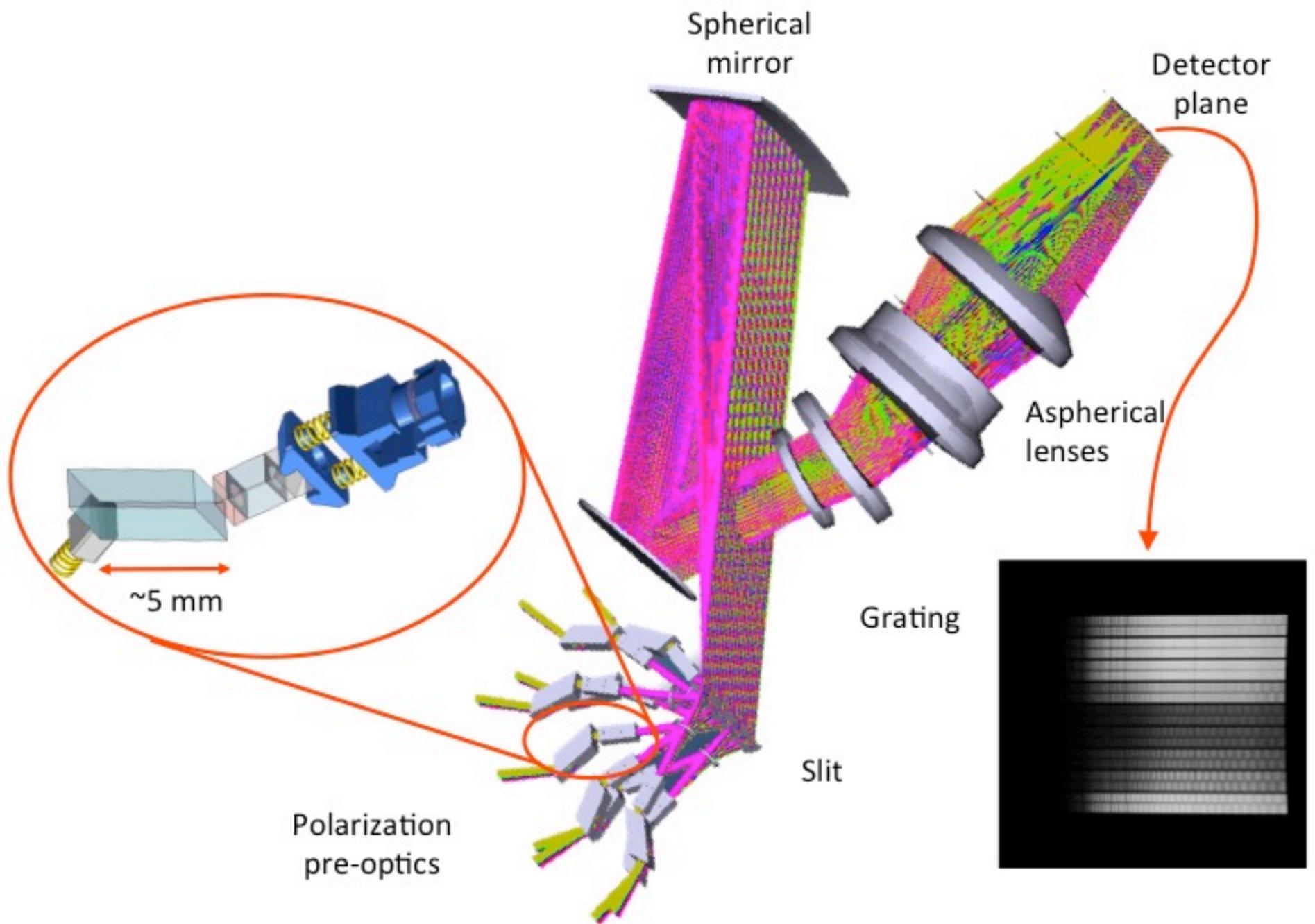
courtesy: Jeroen Rietjens

SPEX measurements @ CESAR, May 2012

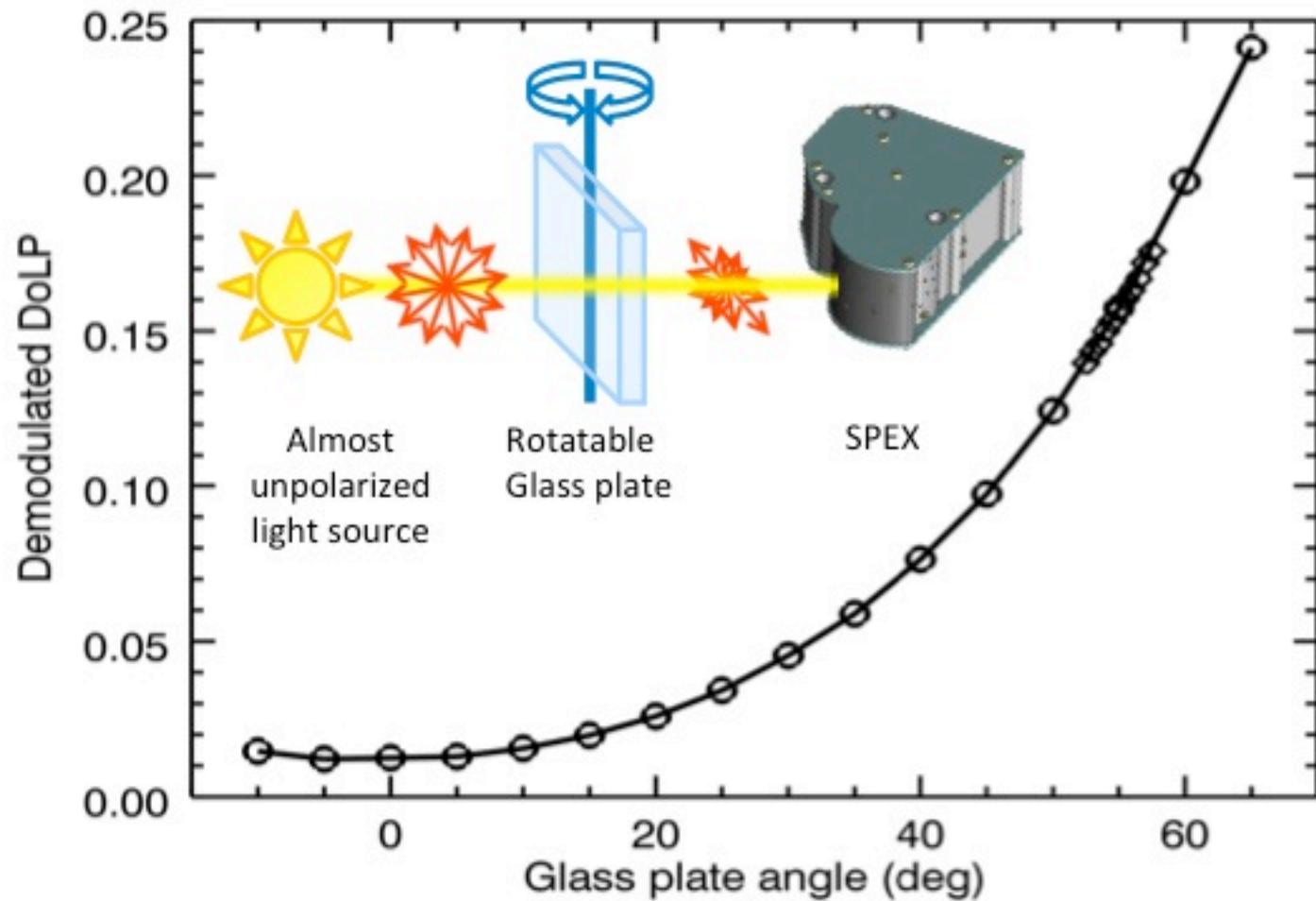


courtesy: Jeroen Rietjens

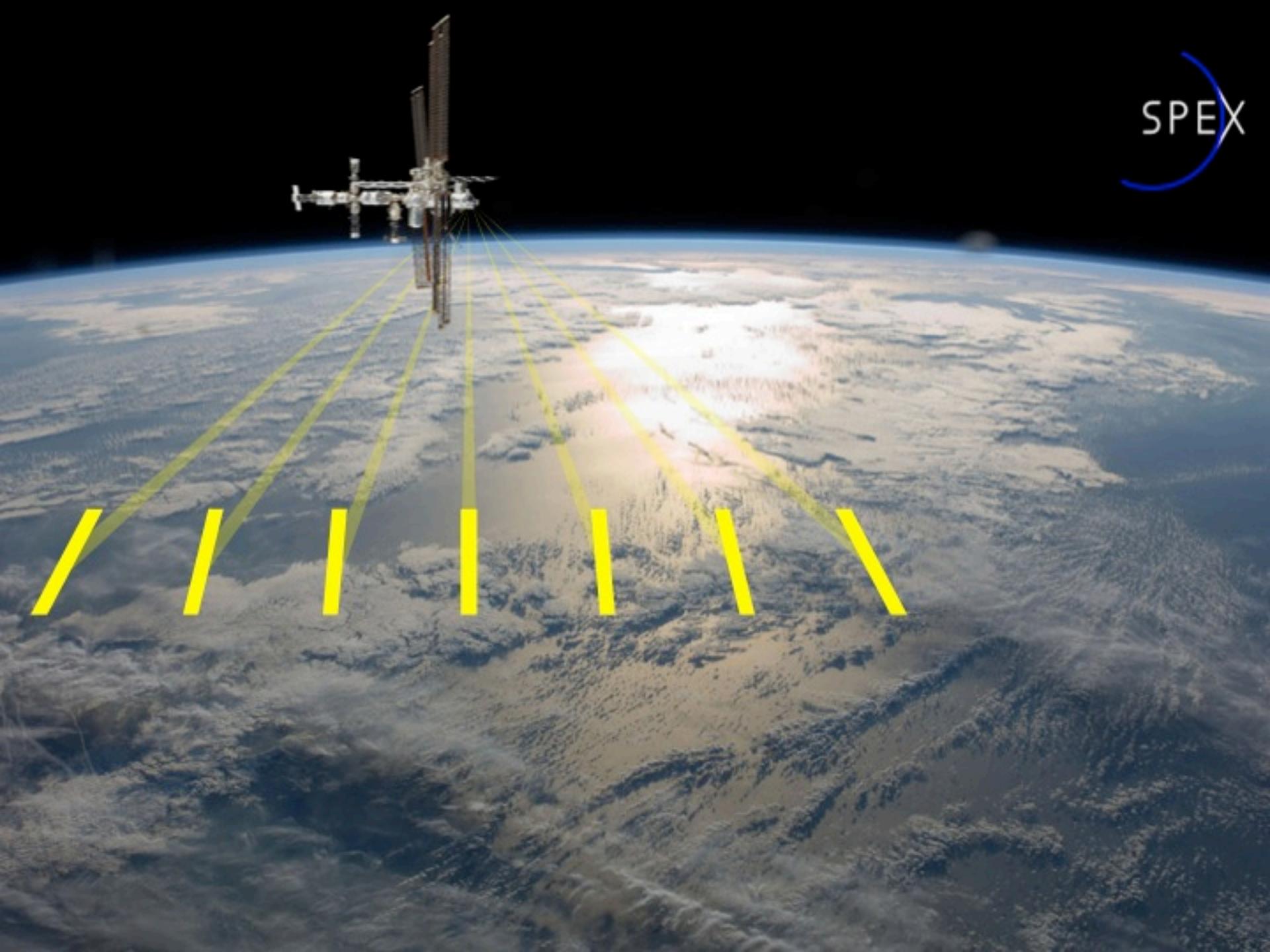




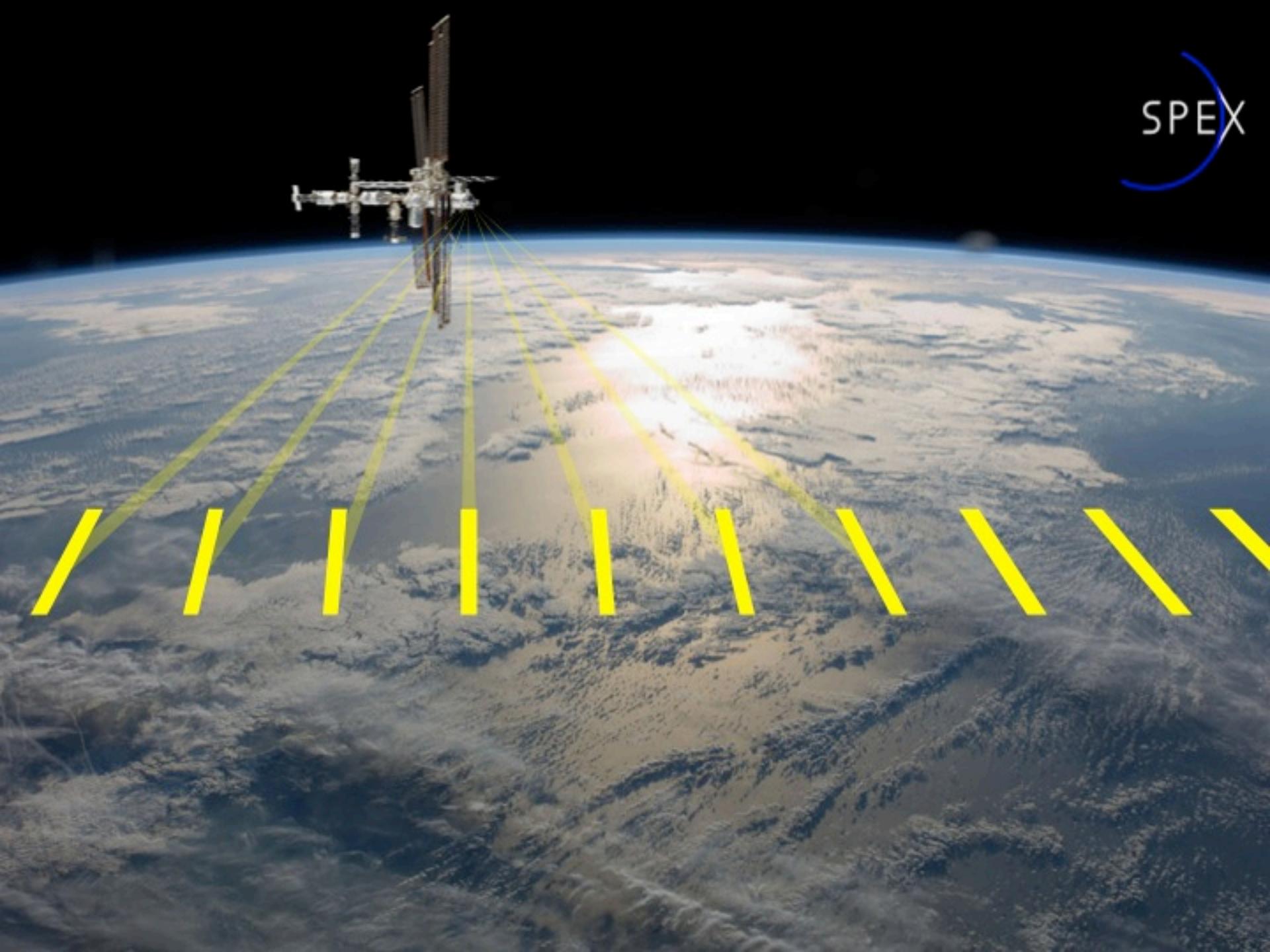
Courtesy of NOVA-ASTRON



- DoLP sensitivity is <0.001
- AoLP sensitivity is <0.2°

A photograph of the International Space Station (ISS) in orbit around Earth. The station is positioned in the upper left quadrant, with its solar panels extended. Several yellow lines descend from the station, pointing towards the Earth's surface below. These lines highlight specific locations on the planet's cloud cover. In the top right corner, the word "SPEX" is written in white capital letters, accompanied by a blue curved arrow that loops around the letter "X".

SPEX

 SPEX

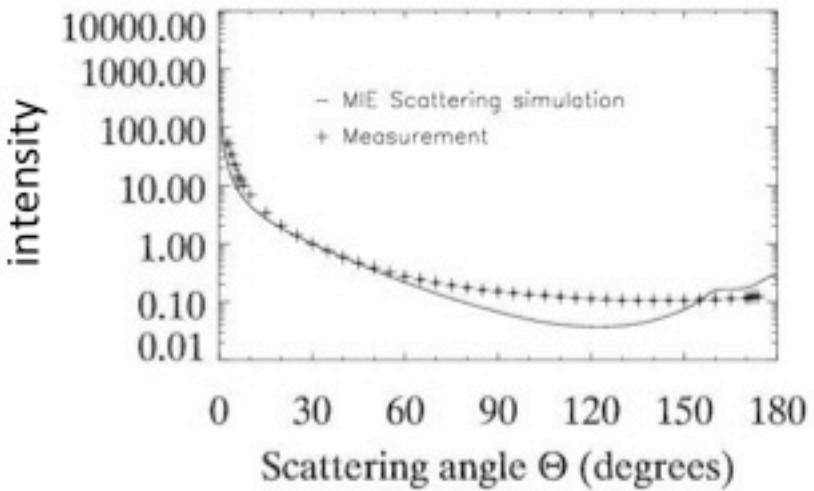
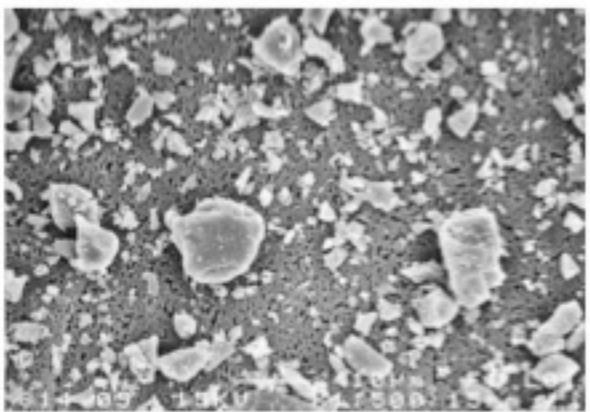


SPEX

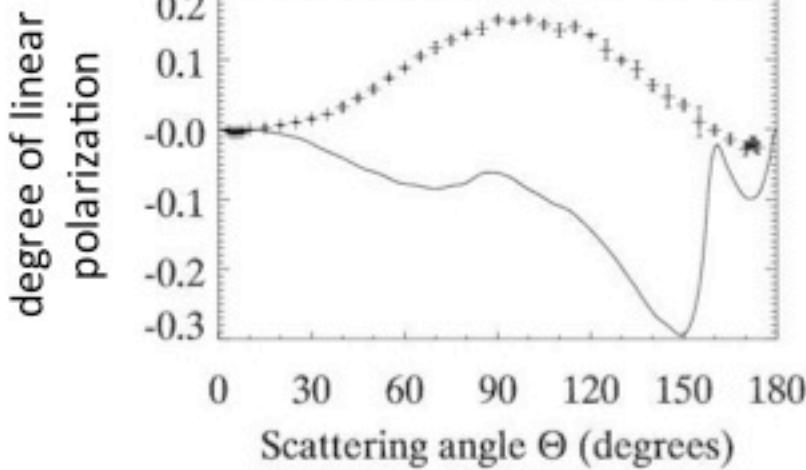
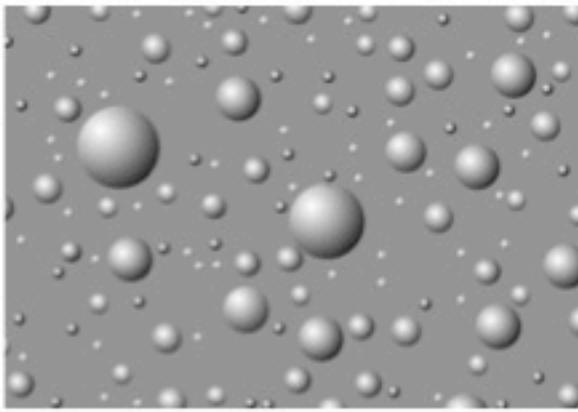




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NOVA

Nederlandse
Onderzoekschool
Voor
Astronomie

SRON

Nederlands
Instituut
voor
Ruimteonderzoek

RIVM

Riks-
Instituut voor
Volksgezondheid en
Milieu

KNMI

Koninklijk
Nederlands
Meteorologisch
Instituut



Measure aerosols with
your smartphone!





prototype



unpolarized sky spectrum



prototype



fluorescent lamp spectrum



prototype



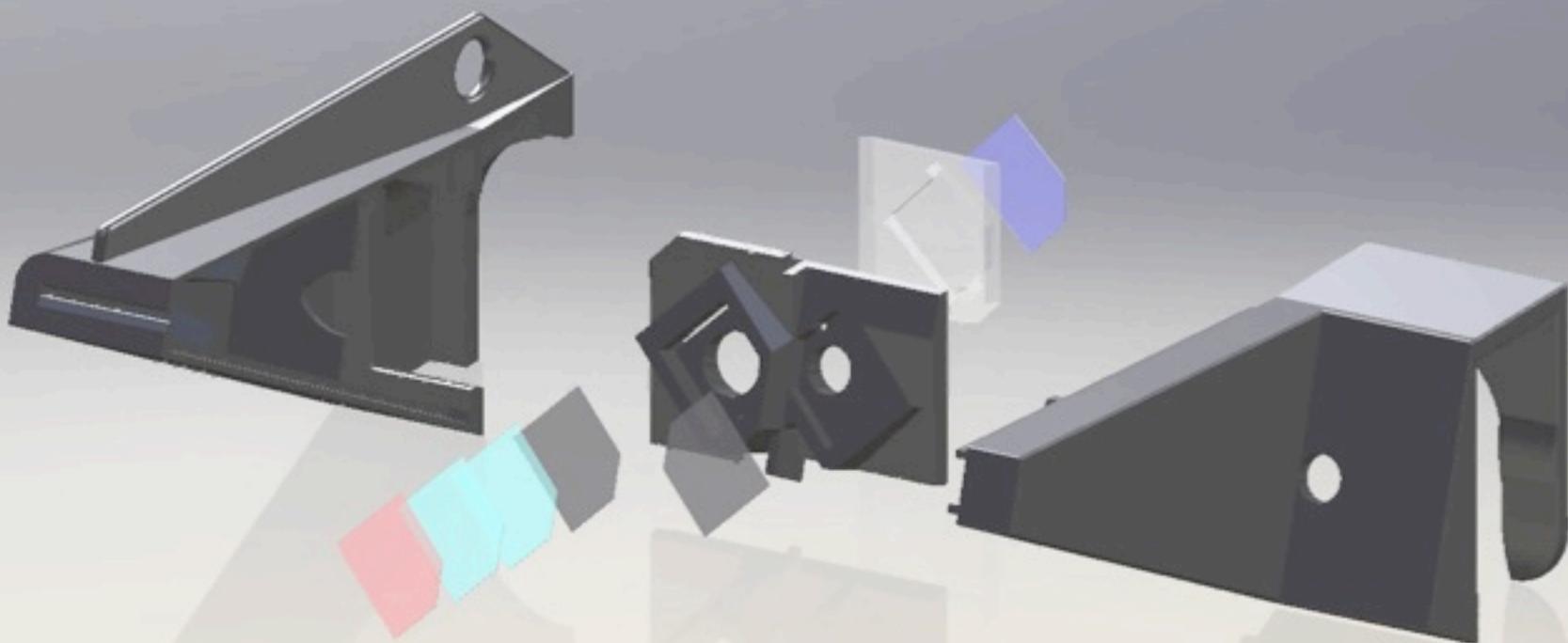
polarized sky spectrum

Winnaar Academische Jaarprijs 2012



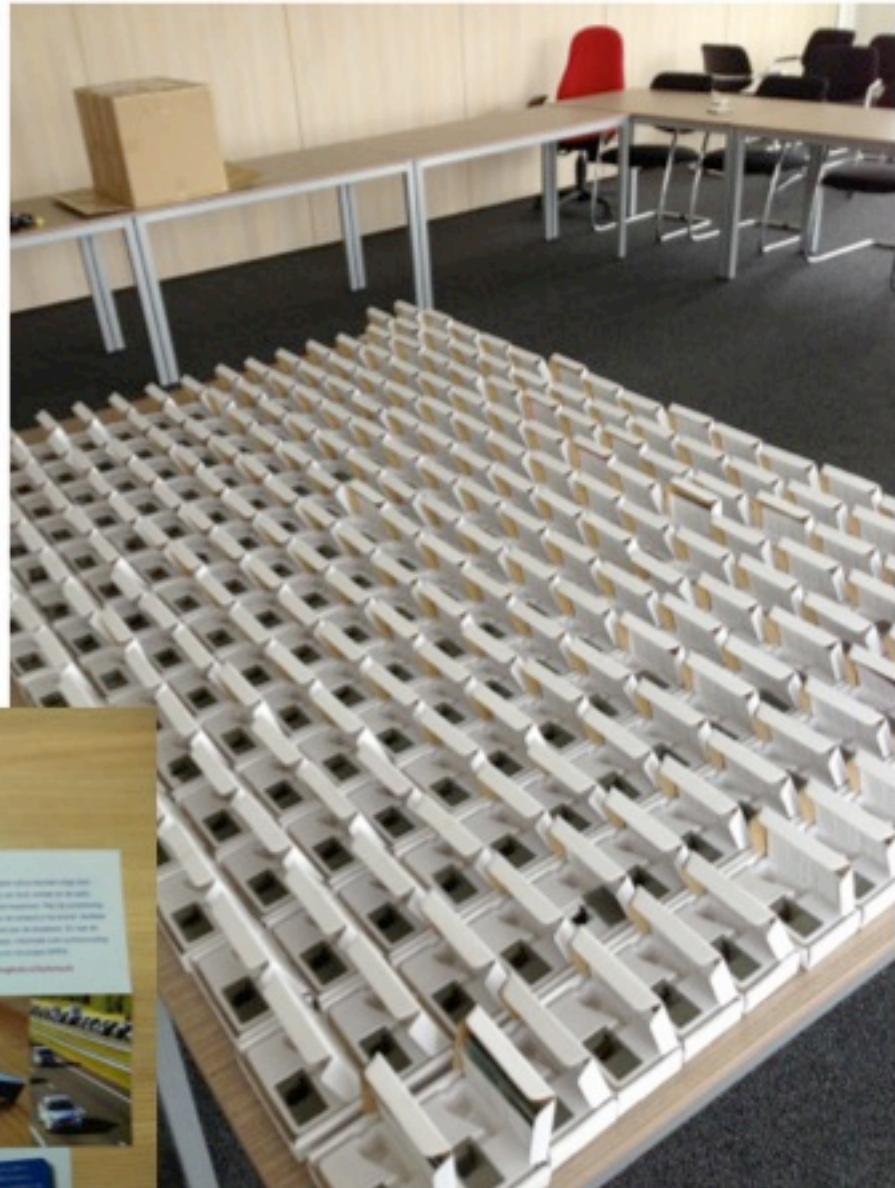


production



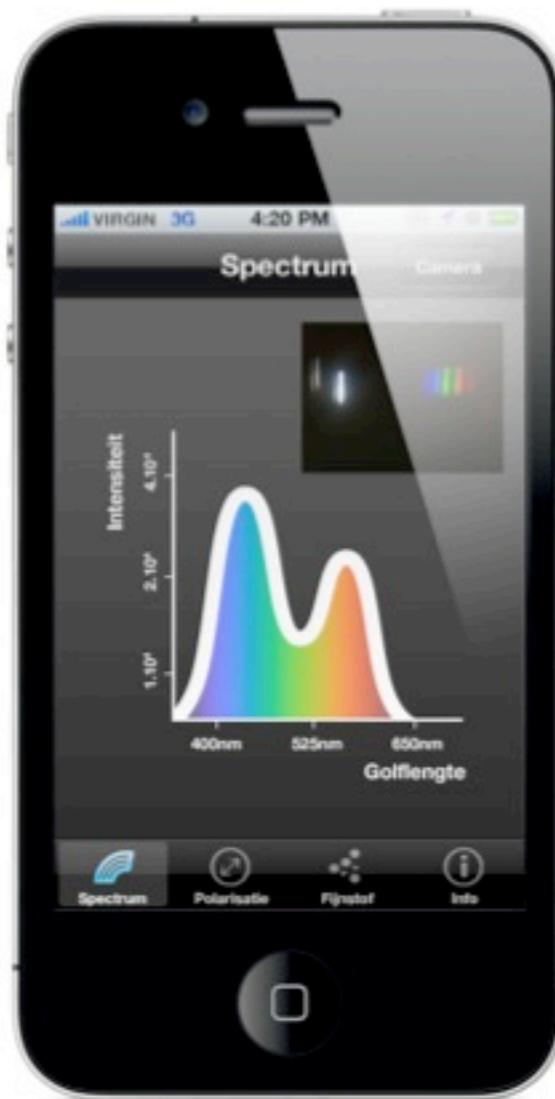


production



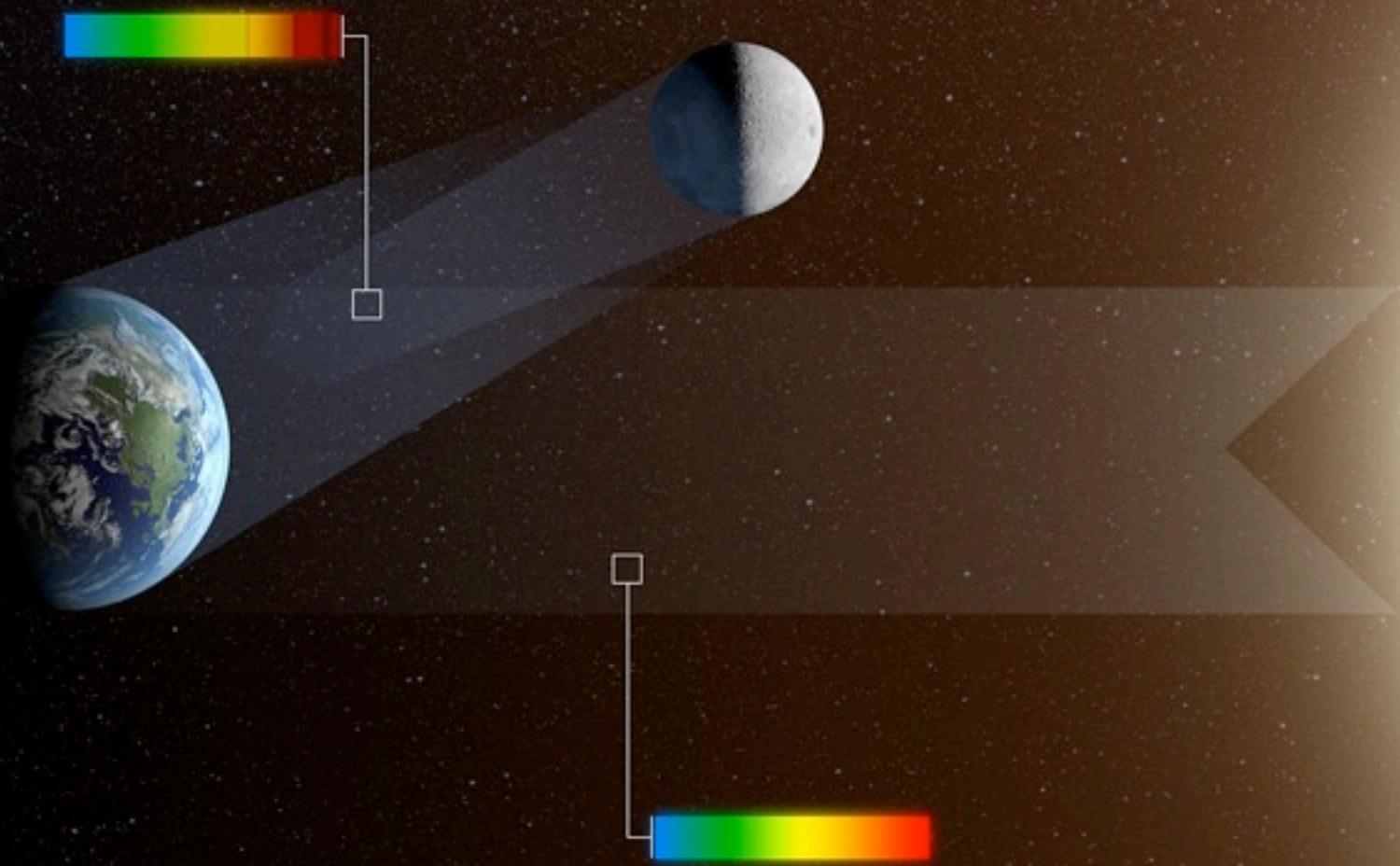


aerosol measurements

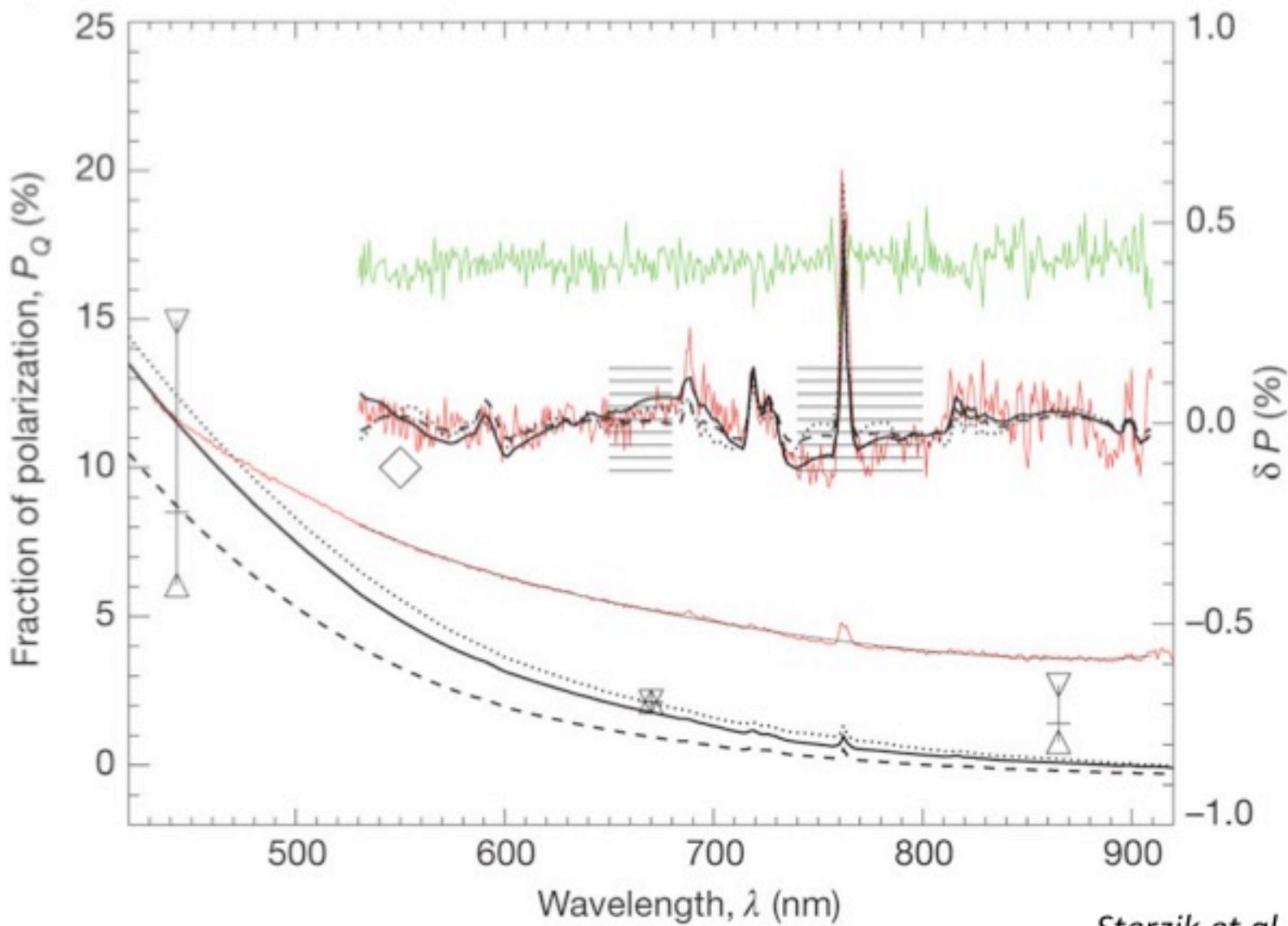




polarization spectrum of the Earthshine



polarization spectrum of the Earthshine

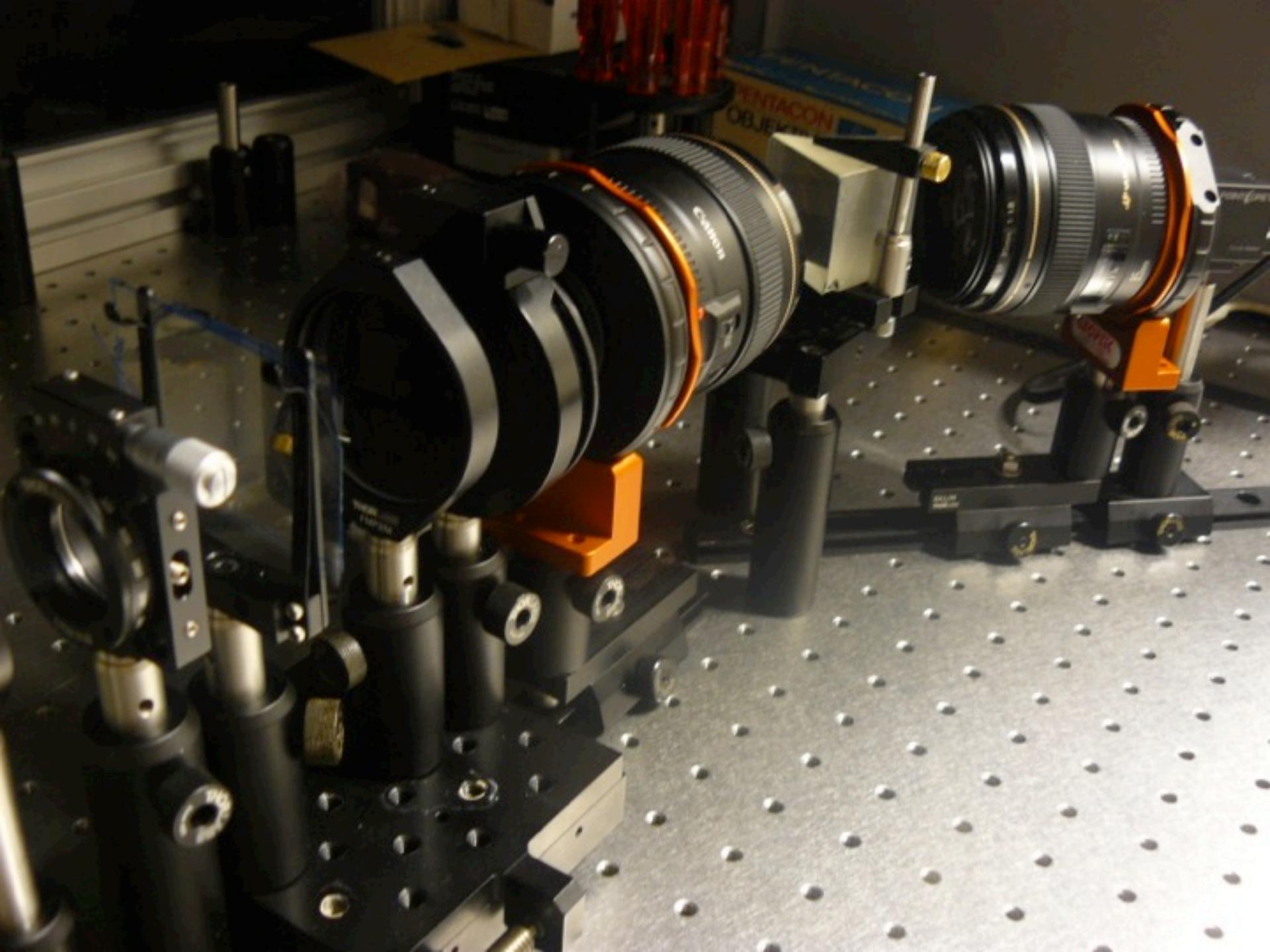


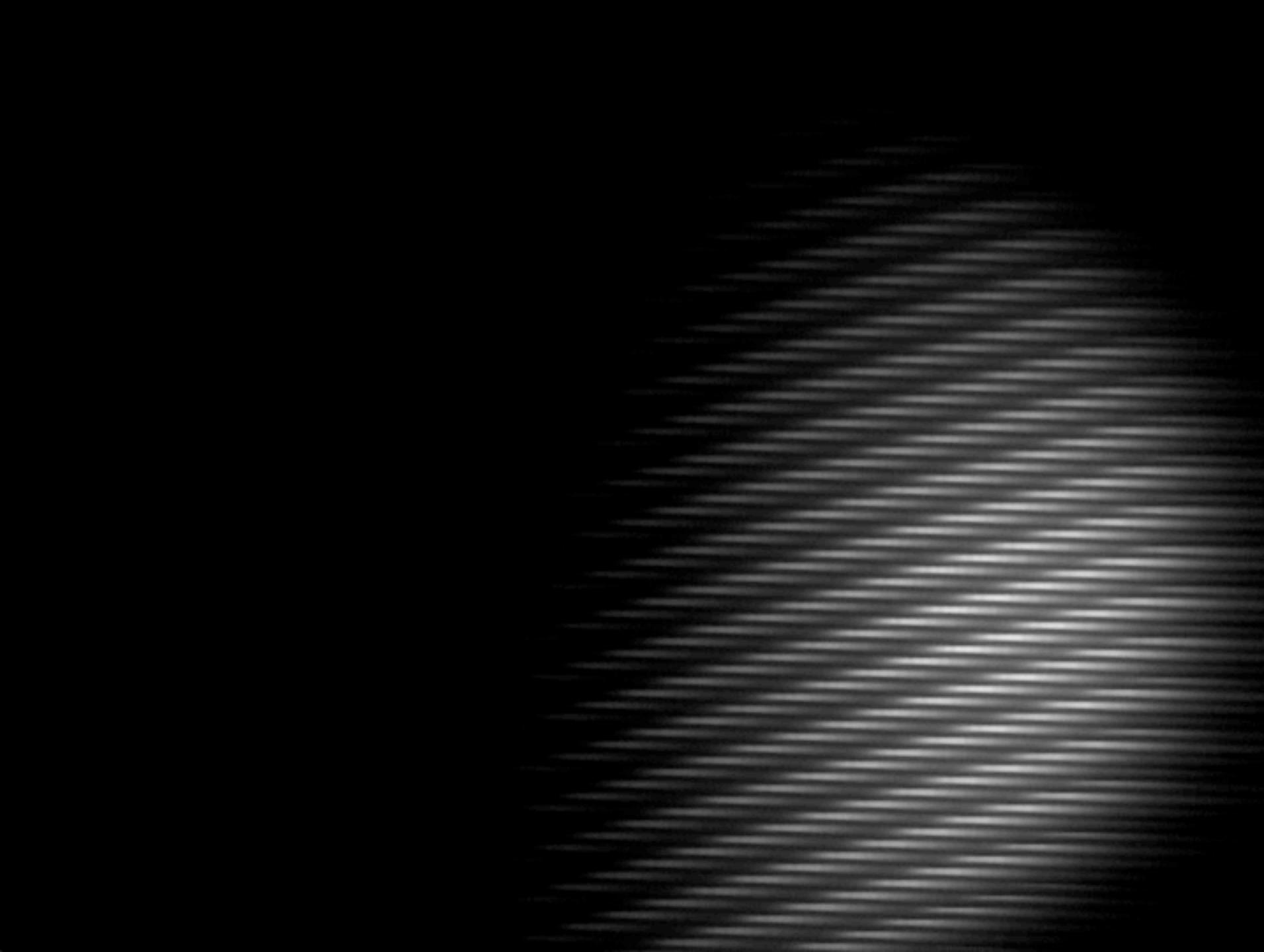
Sterzik et al. (2012)



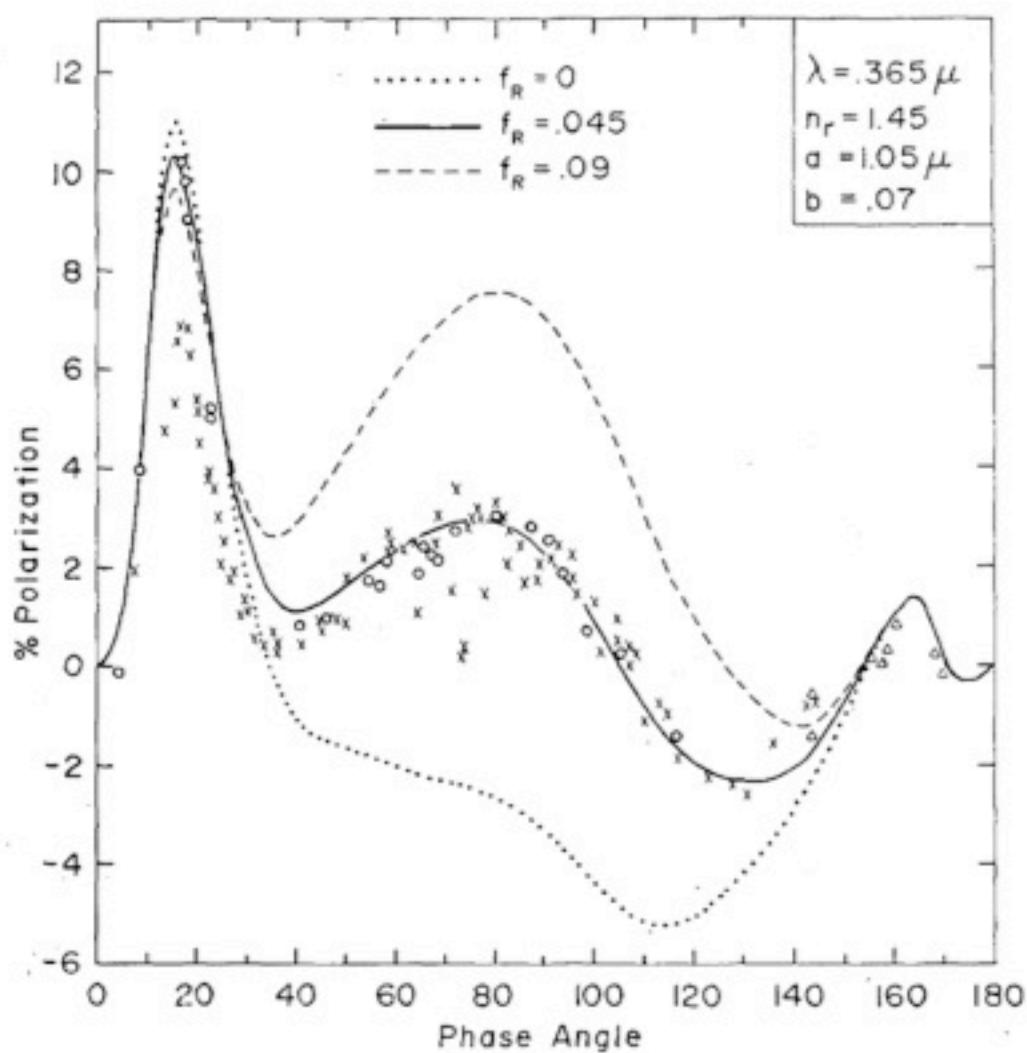
Lunar
Observatory for
Unresolved
Polarimetry of
Earth



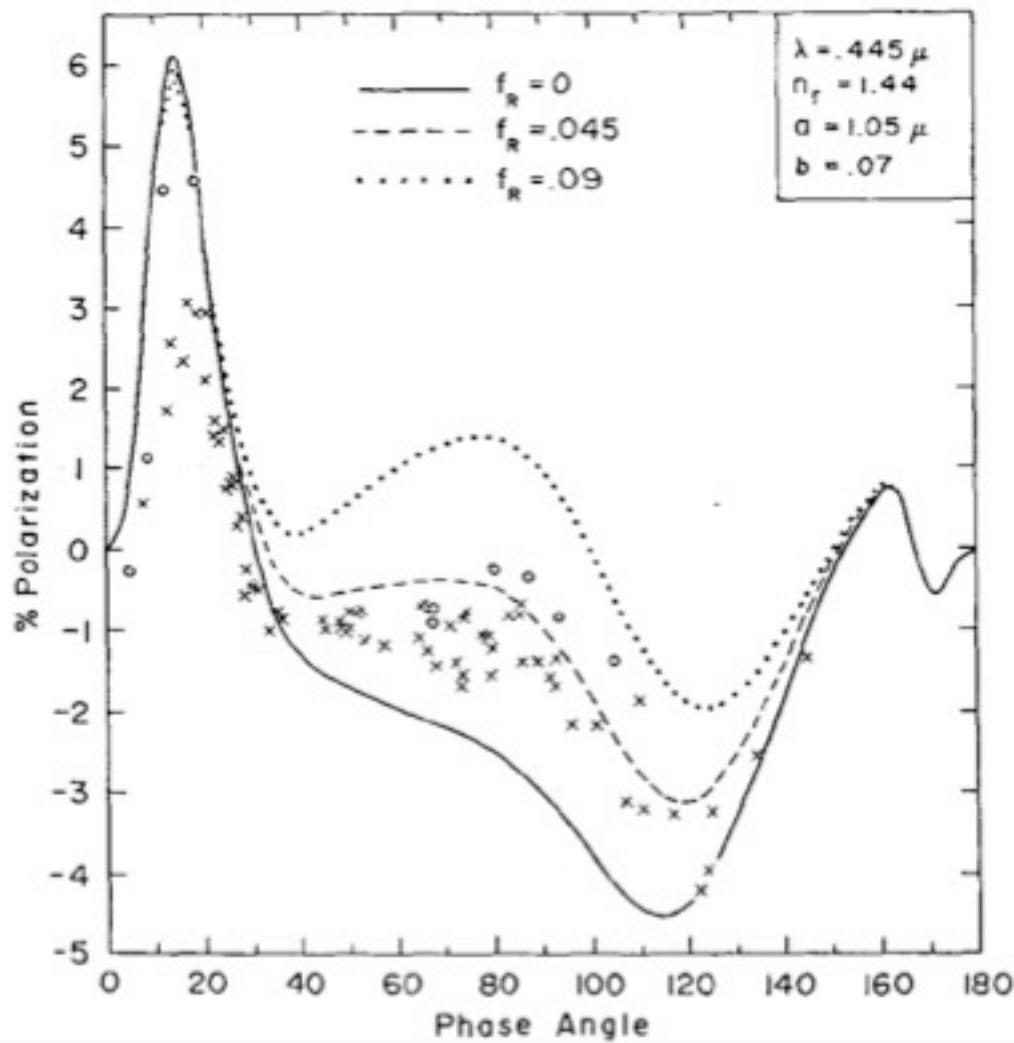




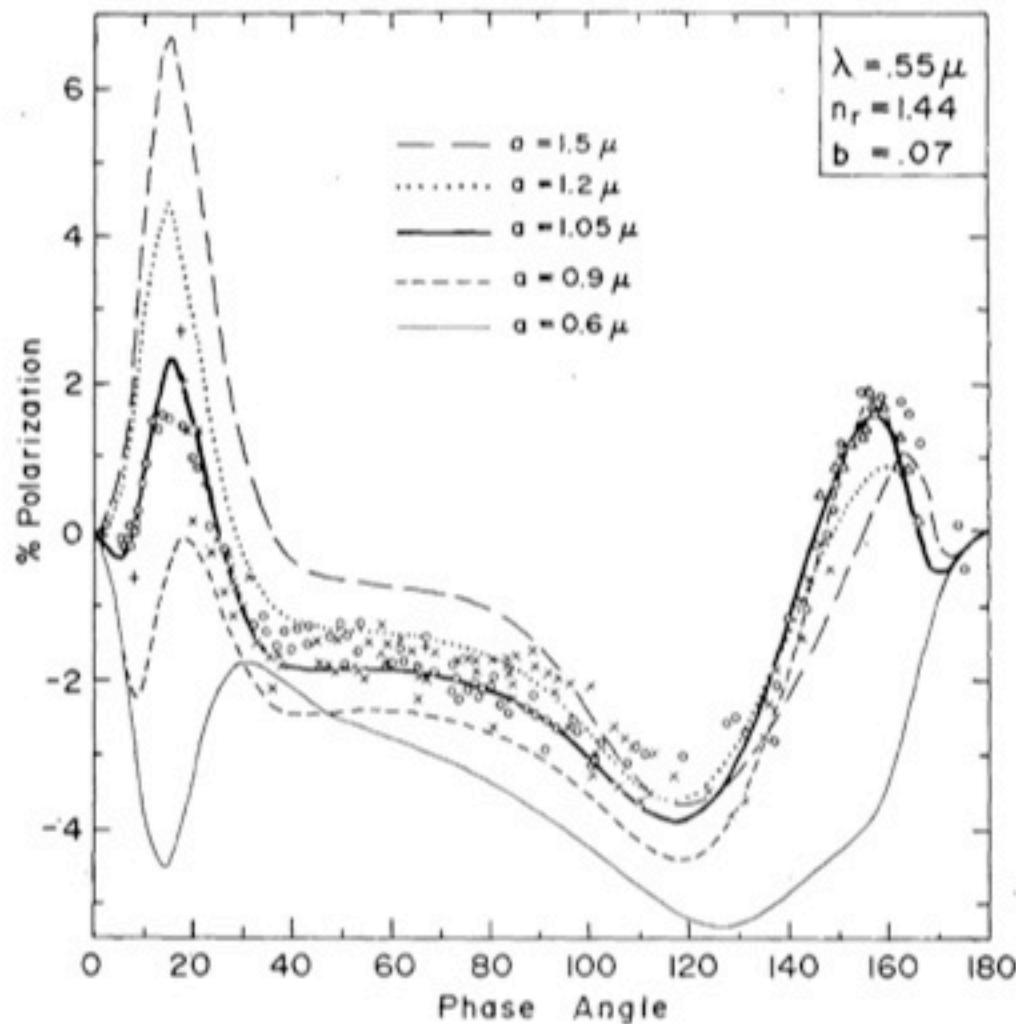
Polarimetry of Venus



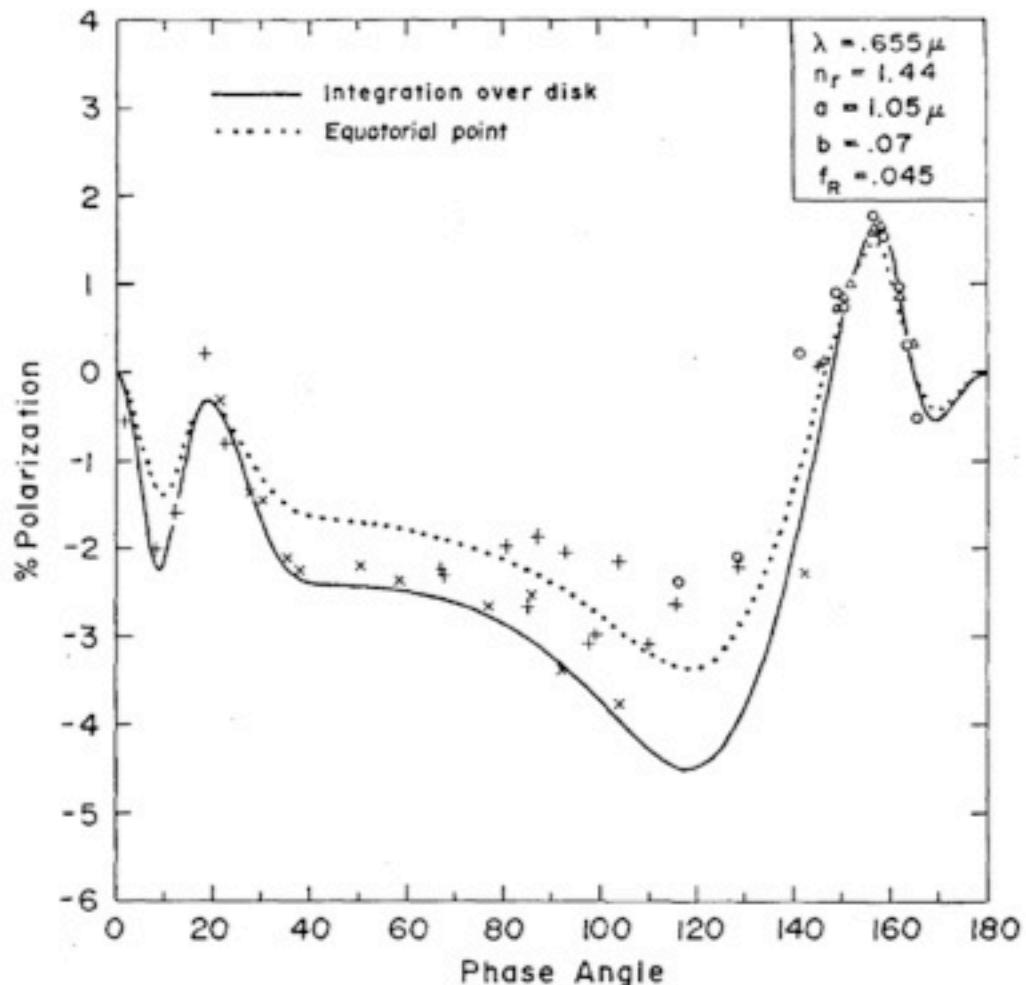
Polarimetry of Venus



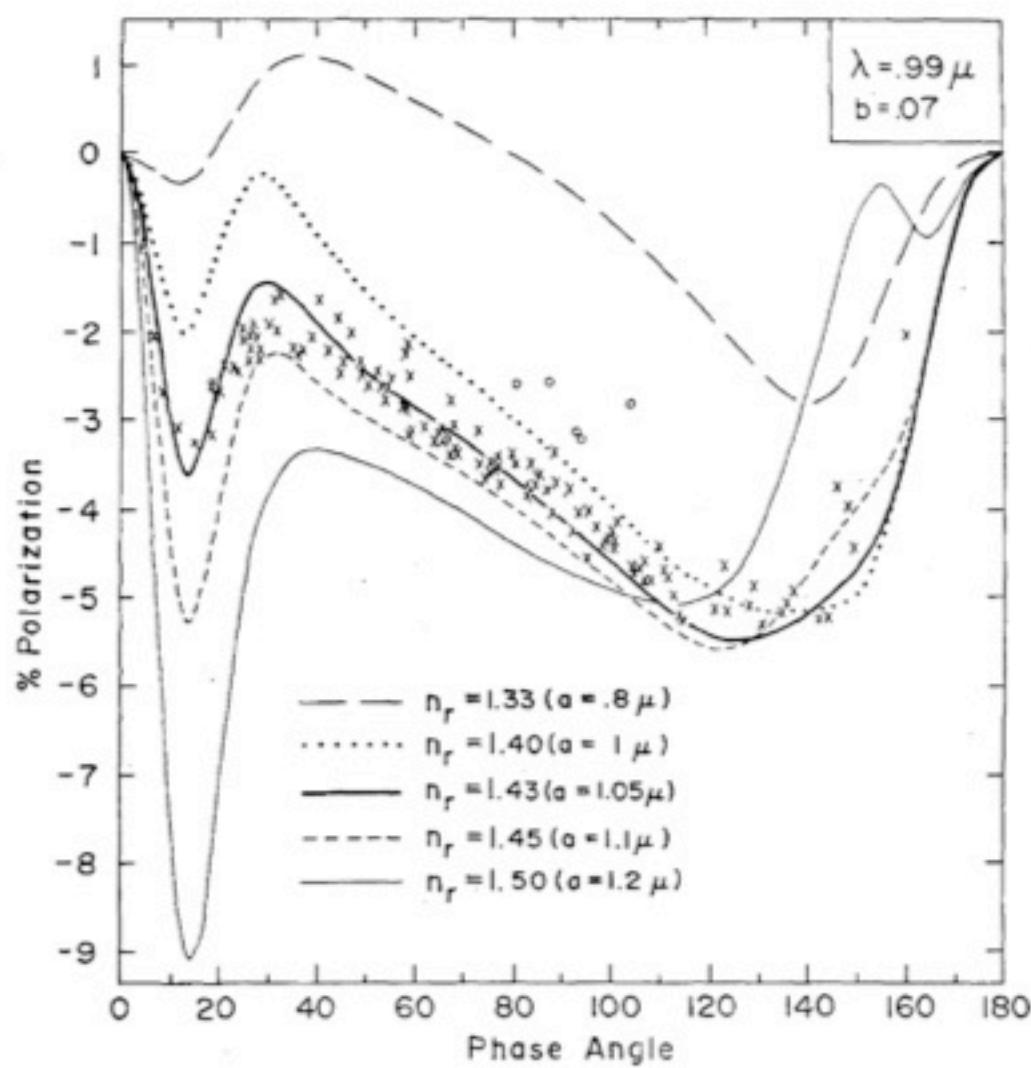
Polarimetry of Venus



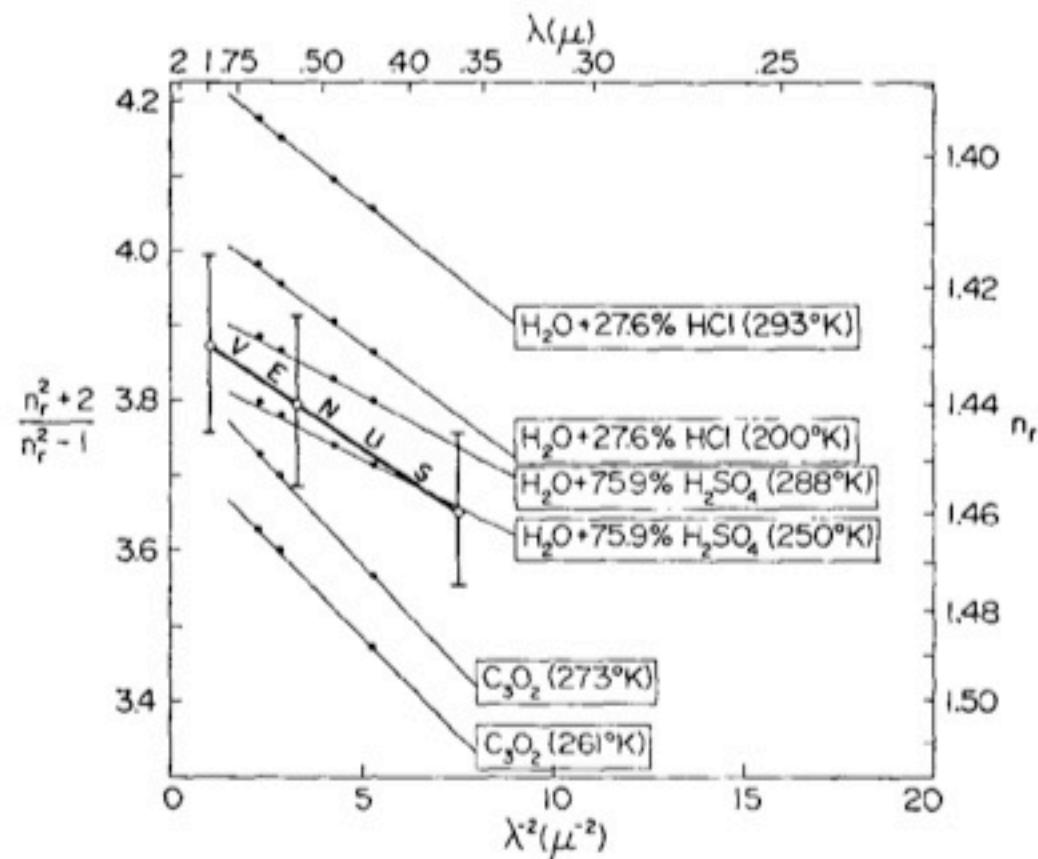
Polarimetry of Venus



Polarimetry of Venus



Polarimetry of Venus

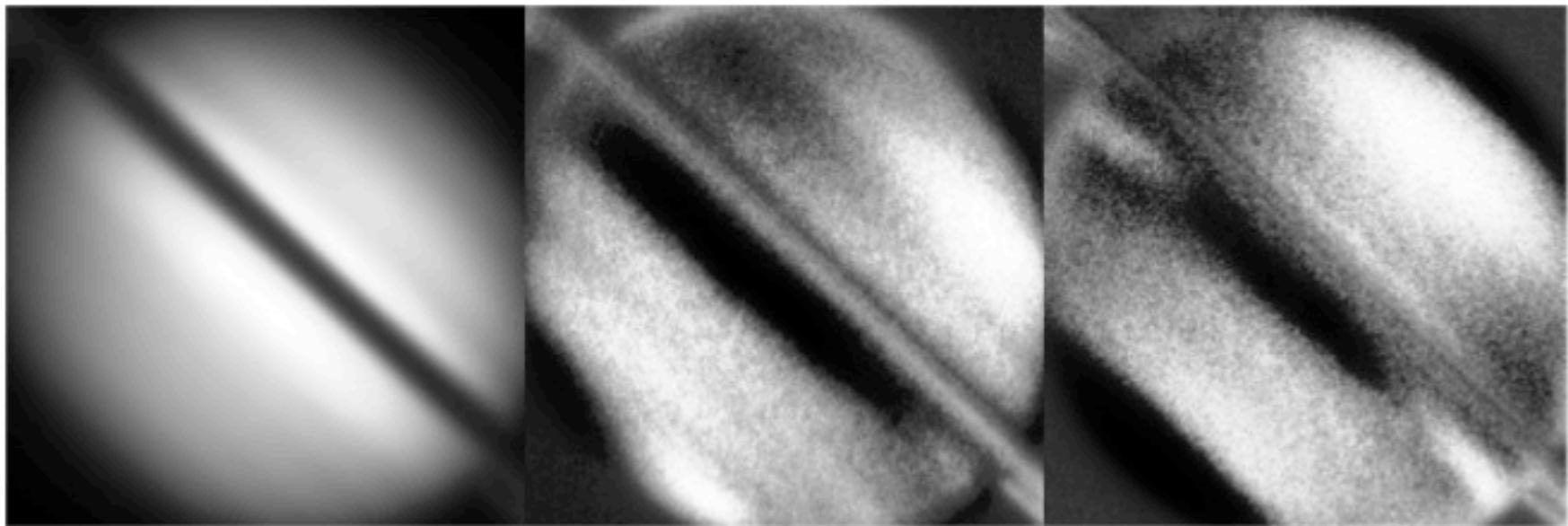


polarized rings around Saturn

I

Q

U

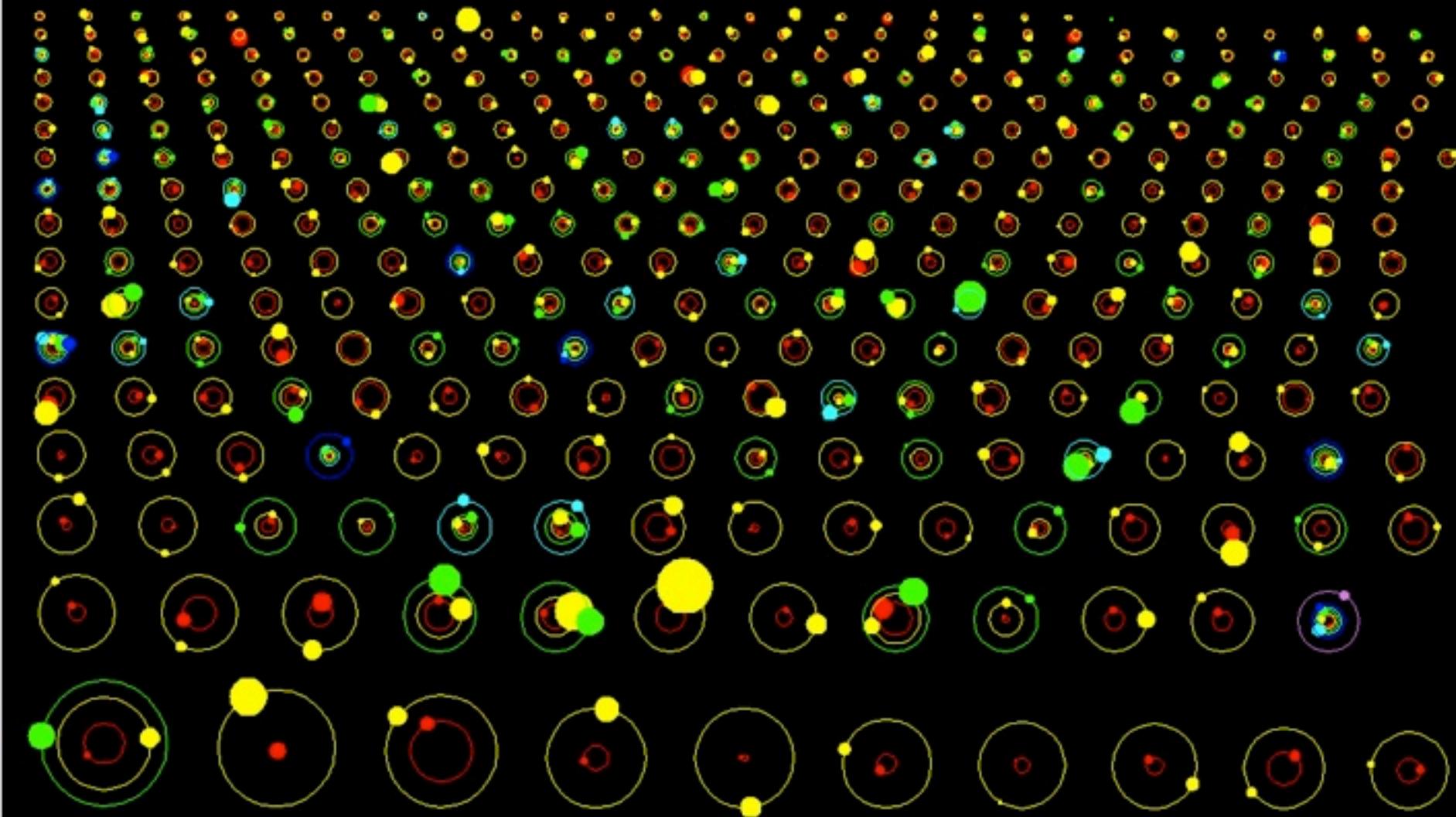


courtesy: Michiel Rodenhuis

The Kepler Orrery II

$t[\text{BJD}] = 2455879$

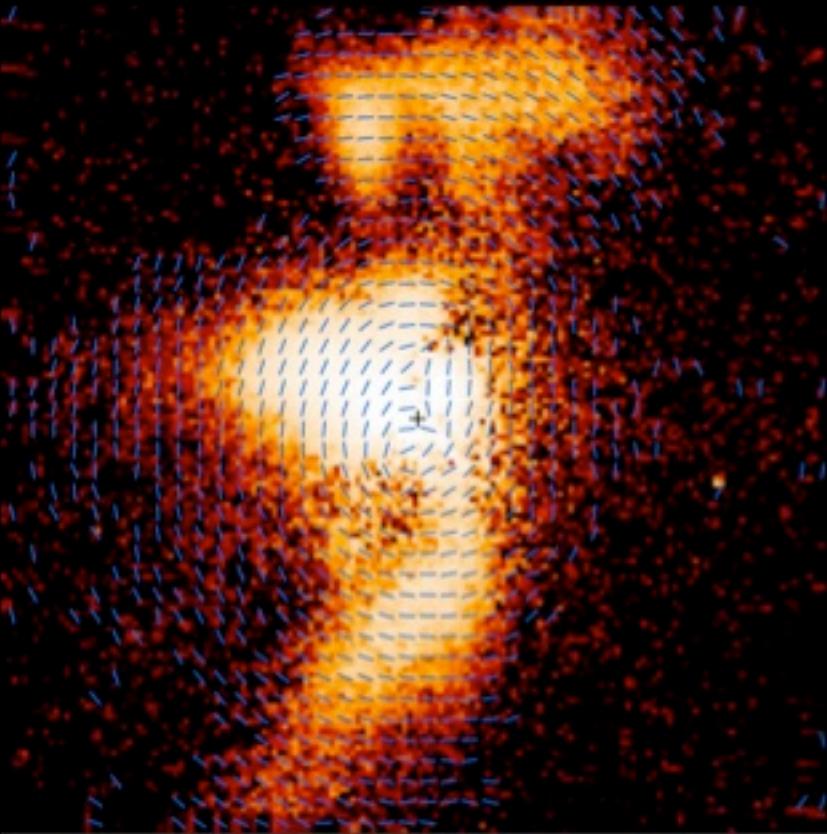
D. Fabrycky 2012



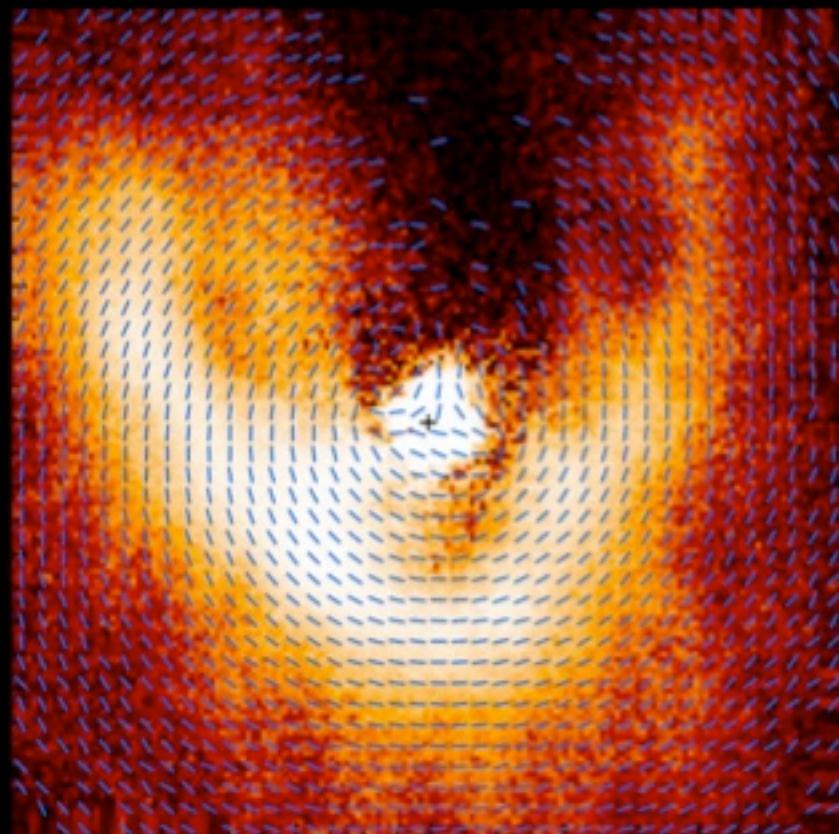


Sirius A&B

ExPo



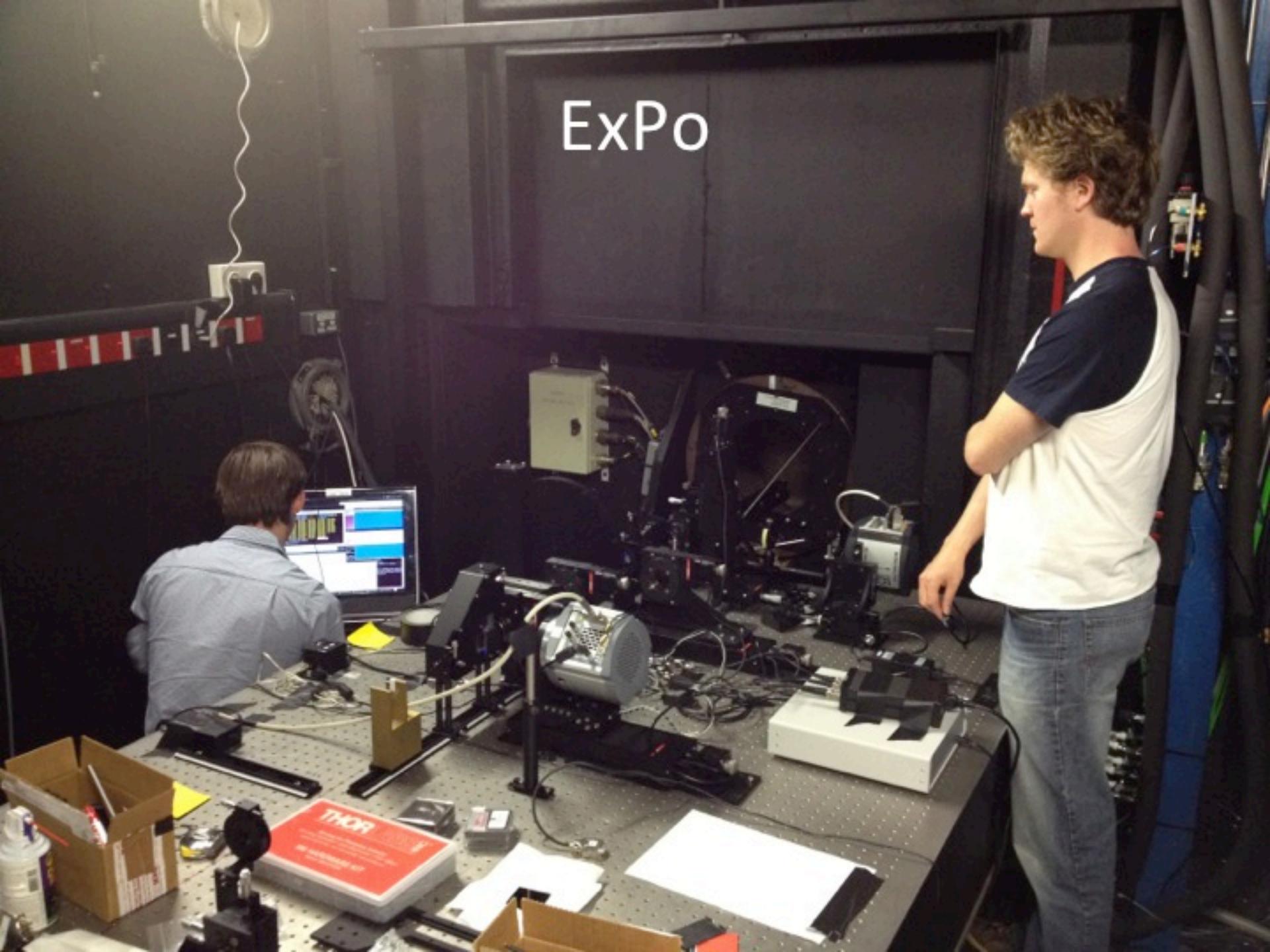
T-Tauri

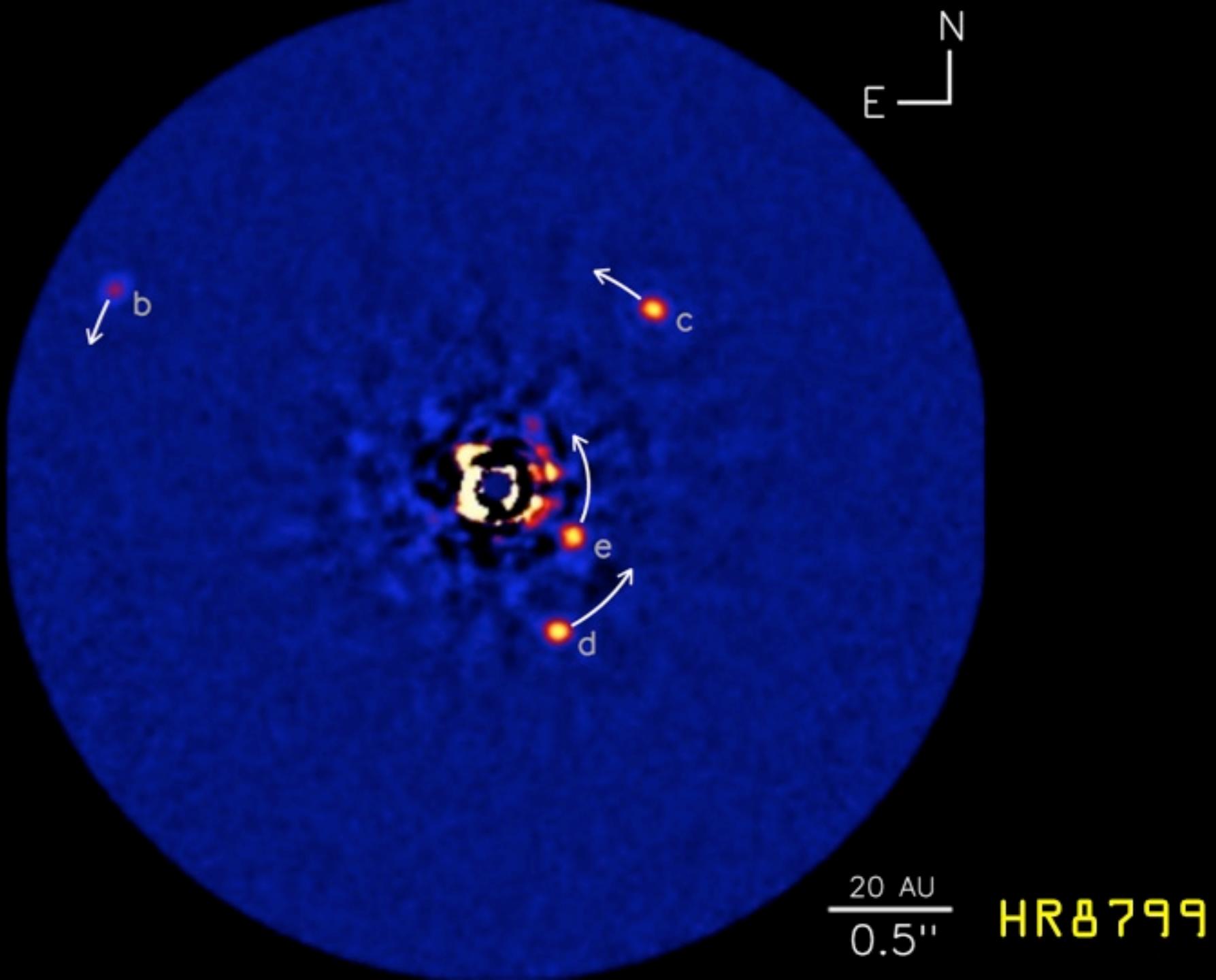


MWC147

courtesy: Michiel Rodenhuis

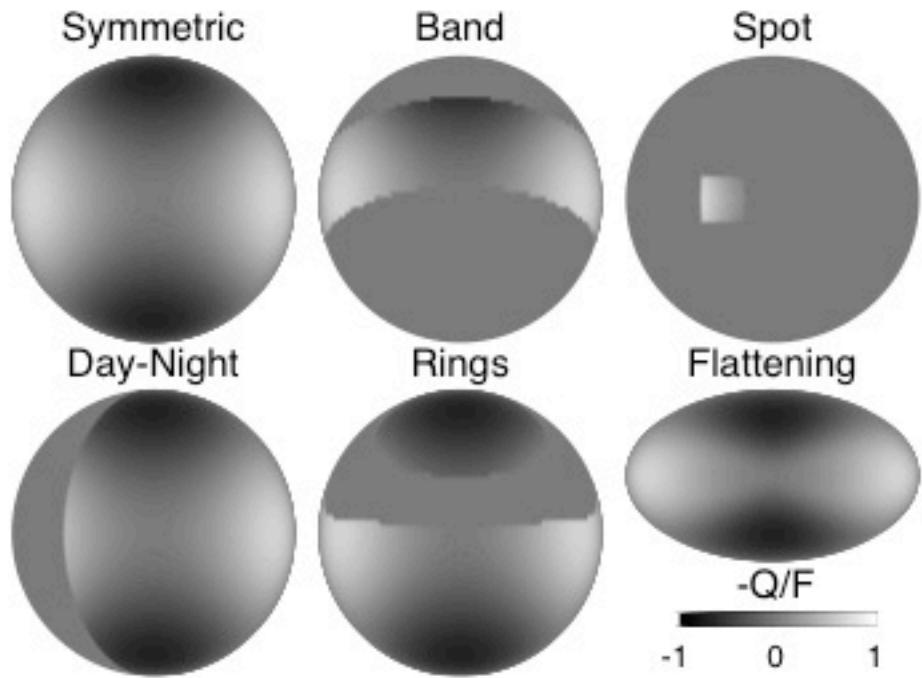
ExPo



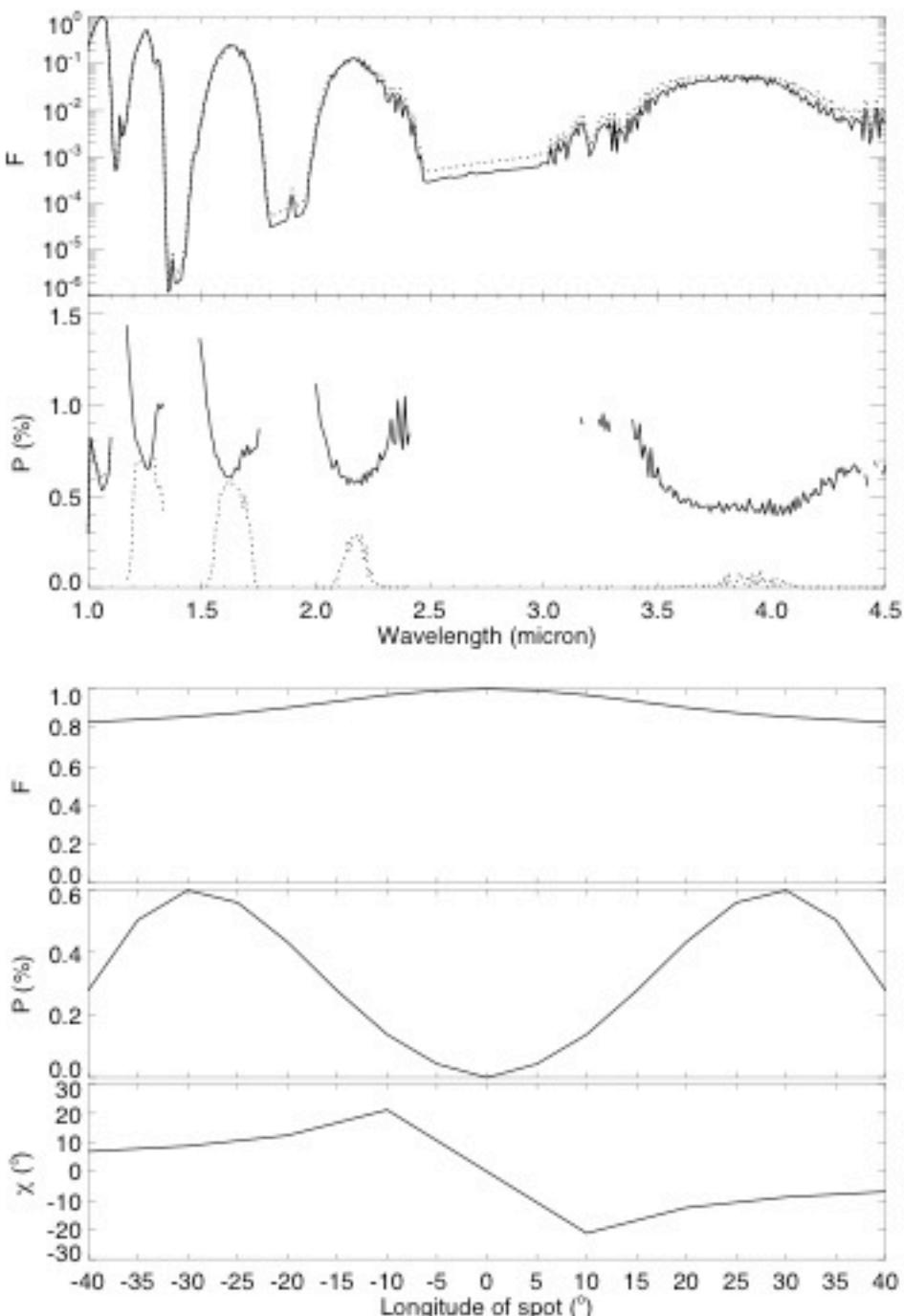


$\frac{20 \text{ AU}}{0.5''}$ HR 8799

The HR 8799 planets do not scatter much star light (too far away). But they do scatter their own thermal emission!



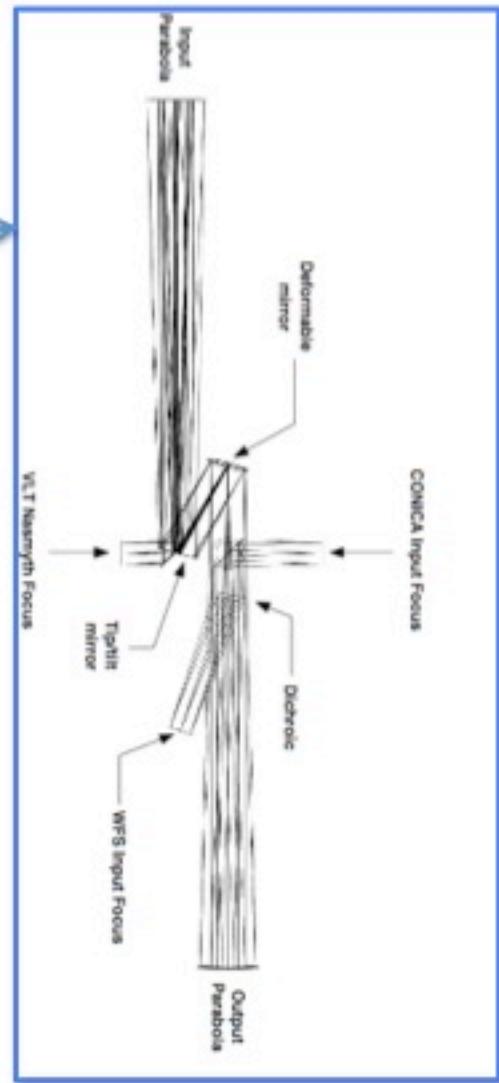
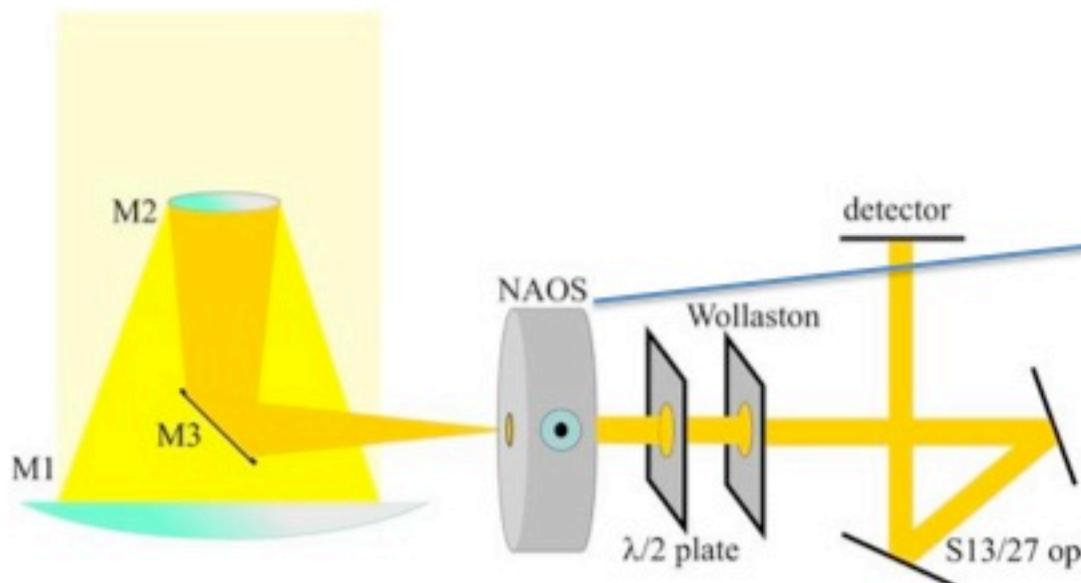
de Kok et al (2011)



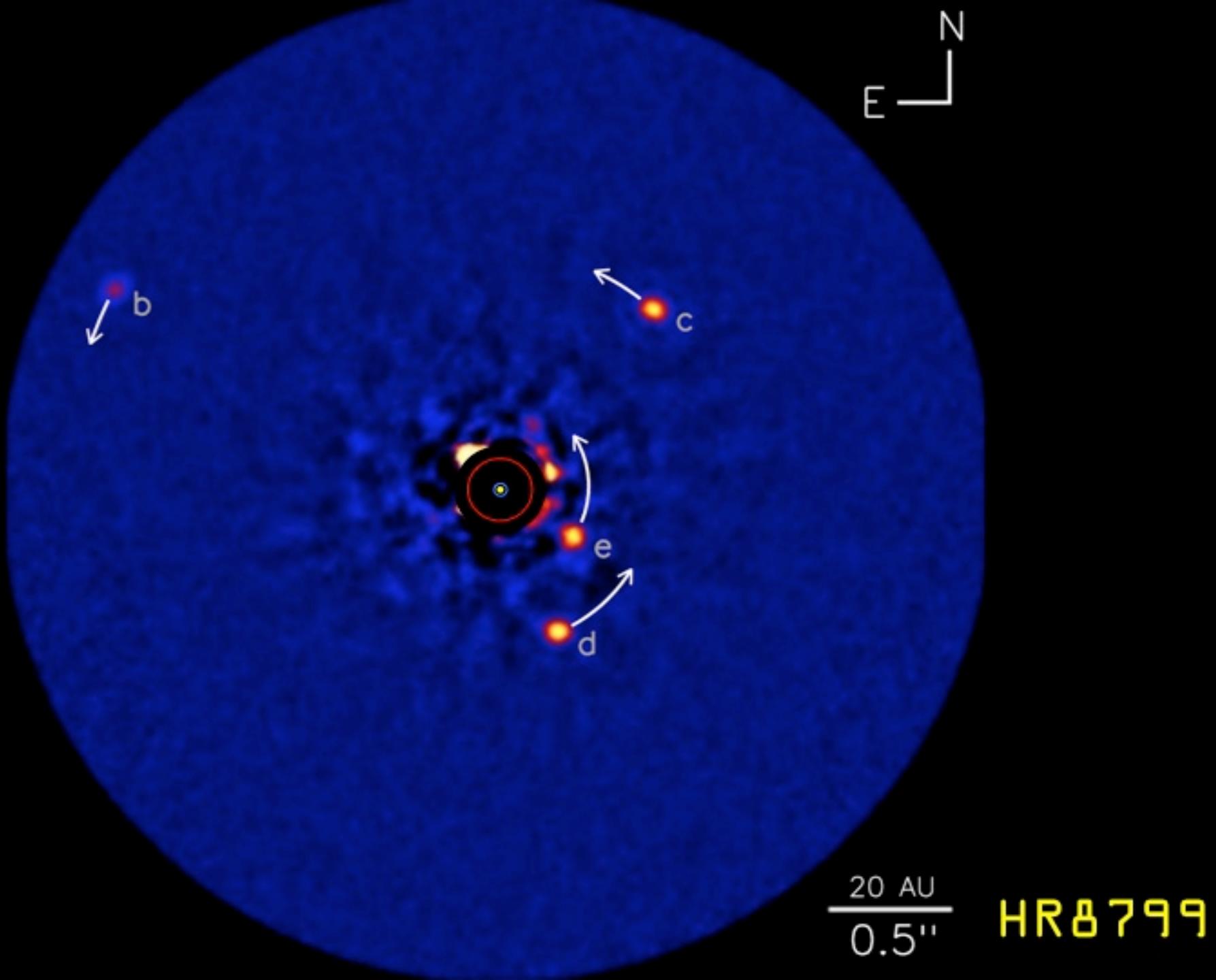
VLT/NaCo



pupil-stabilized polarimetry

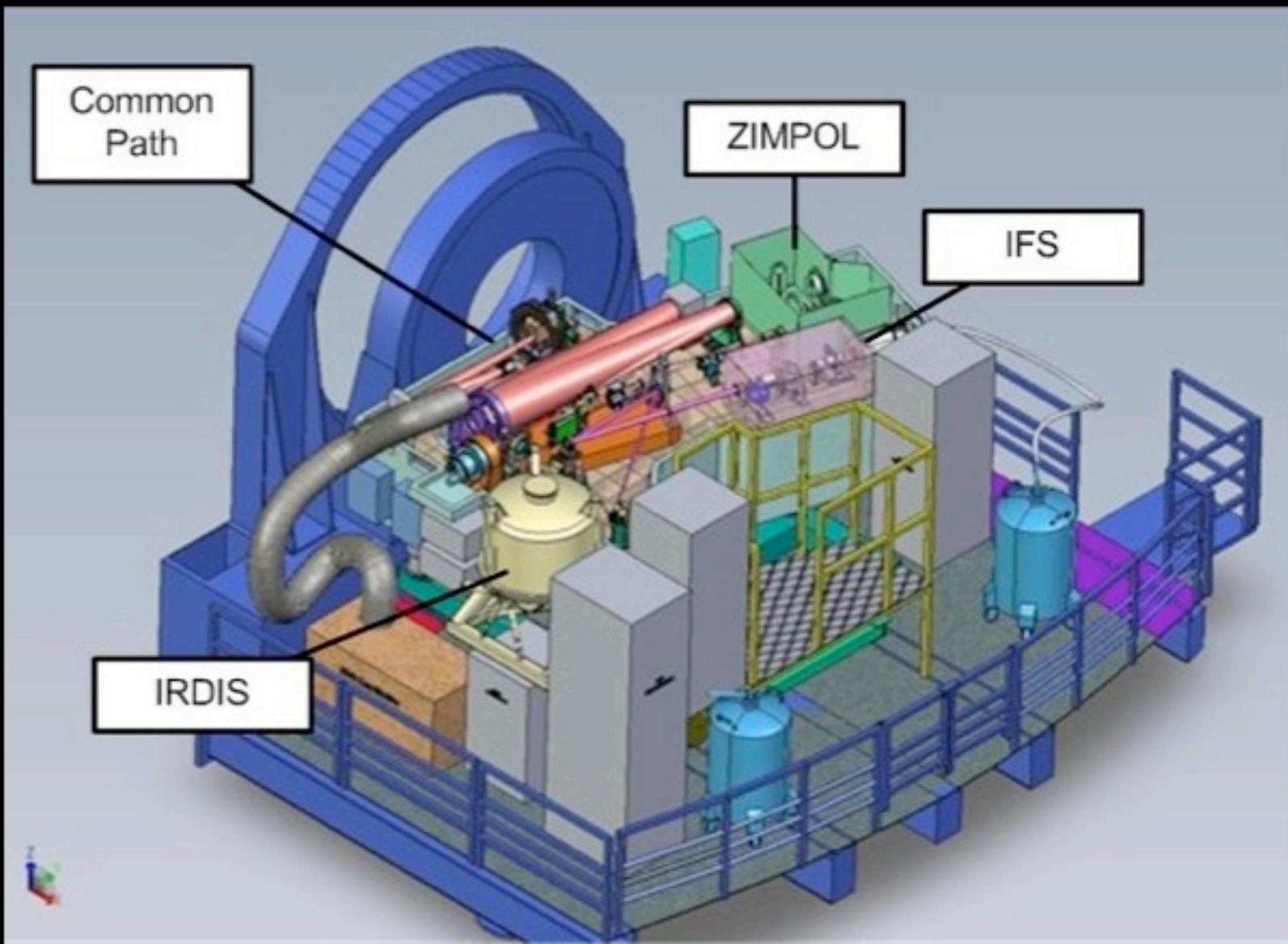


- corotate NaCo with telescope
- IP stabilized & minimized
 - calibrated on standard stars
- field rotates w.r.t speckles
 - PCA PSF subtraction

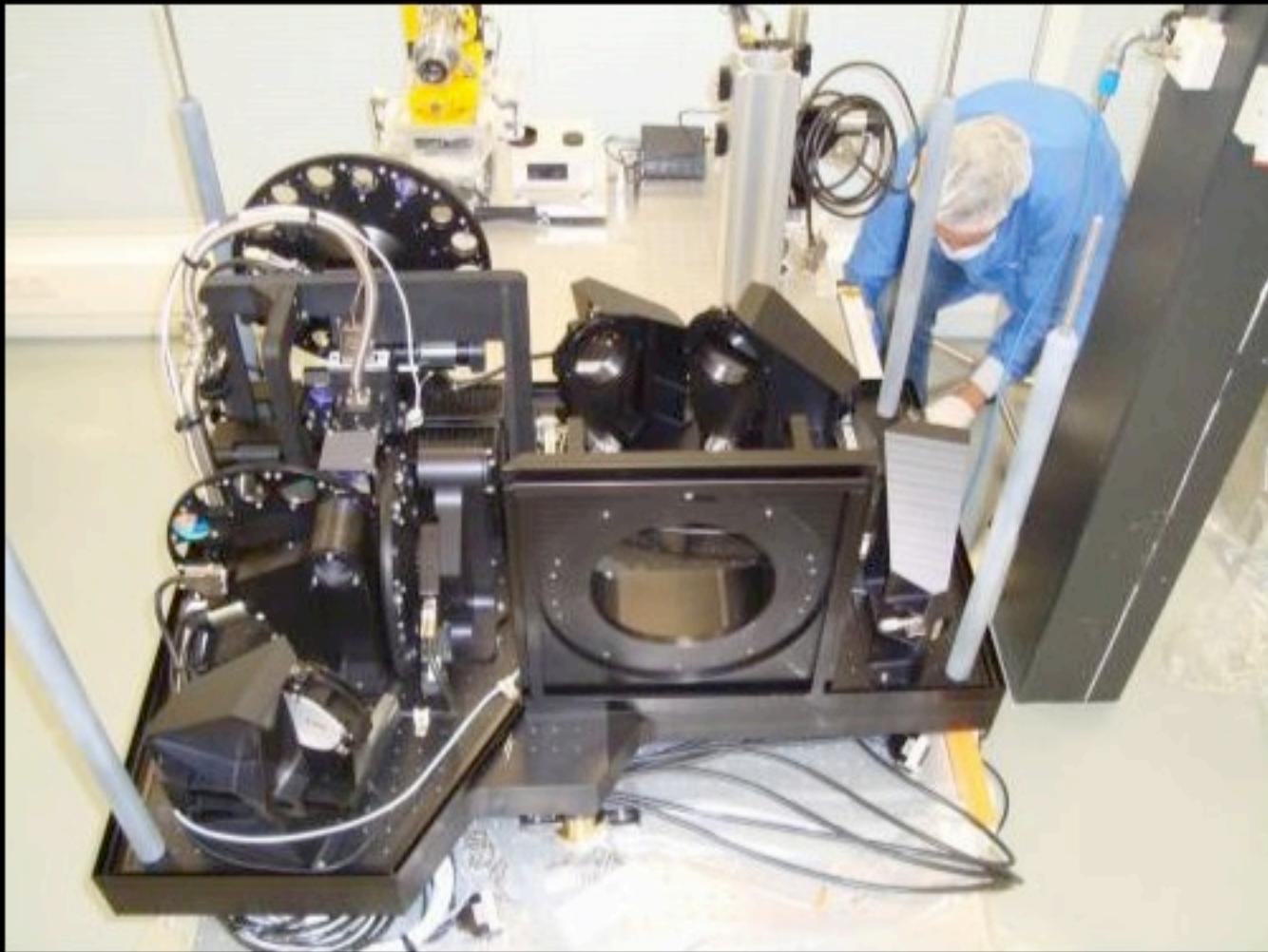




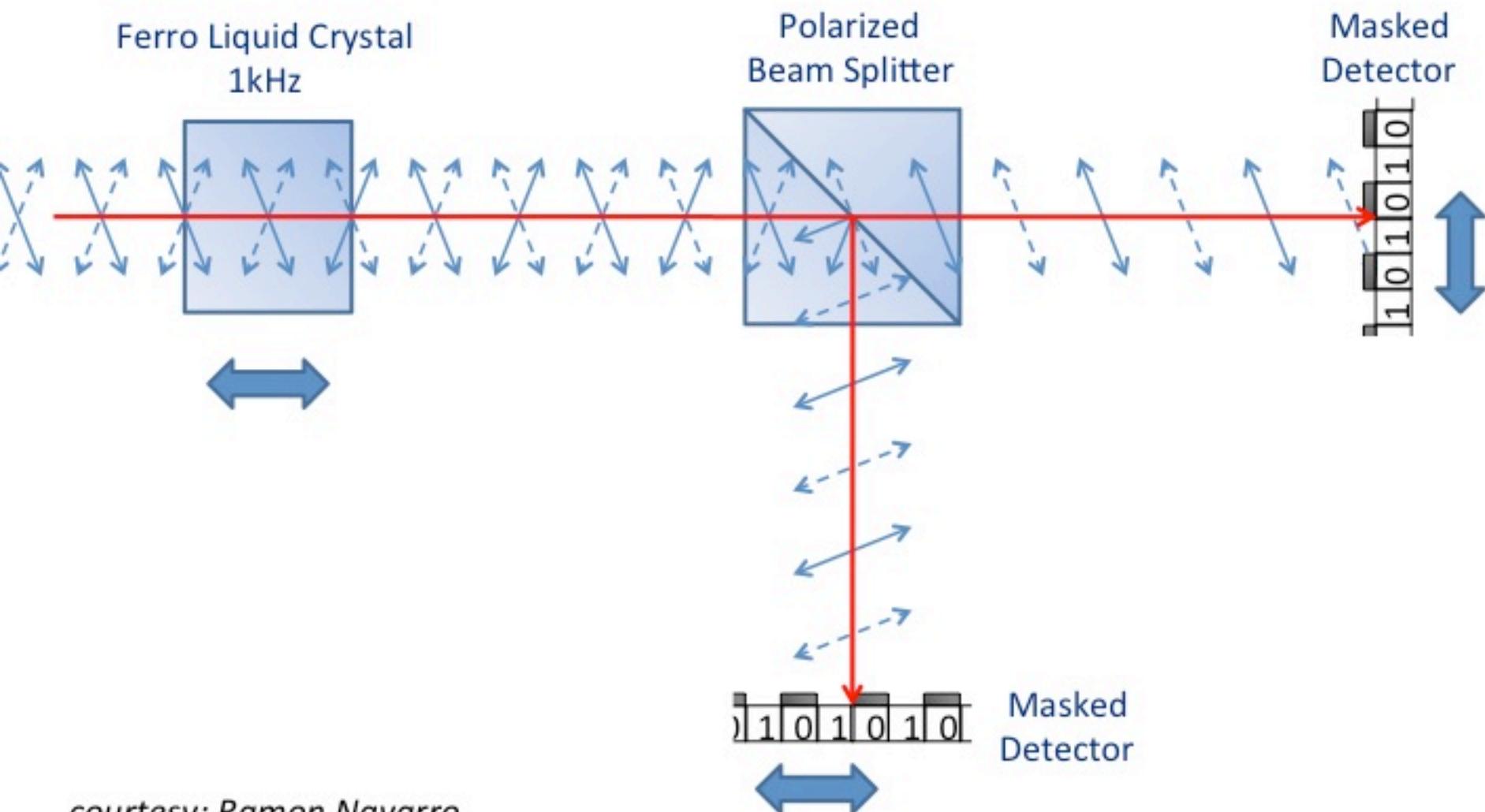
VLT/SPHERE



VLT/SPHERE-ZIMPOL

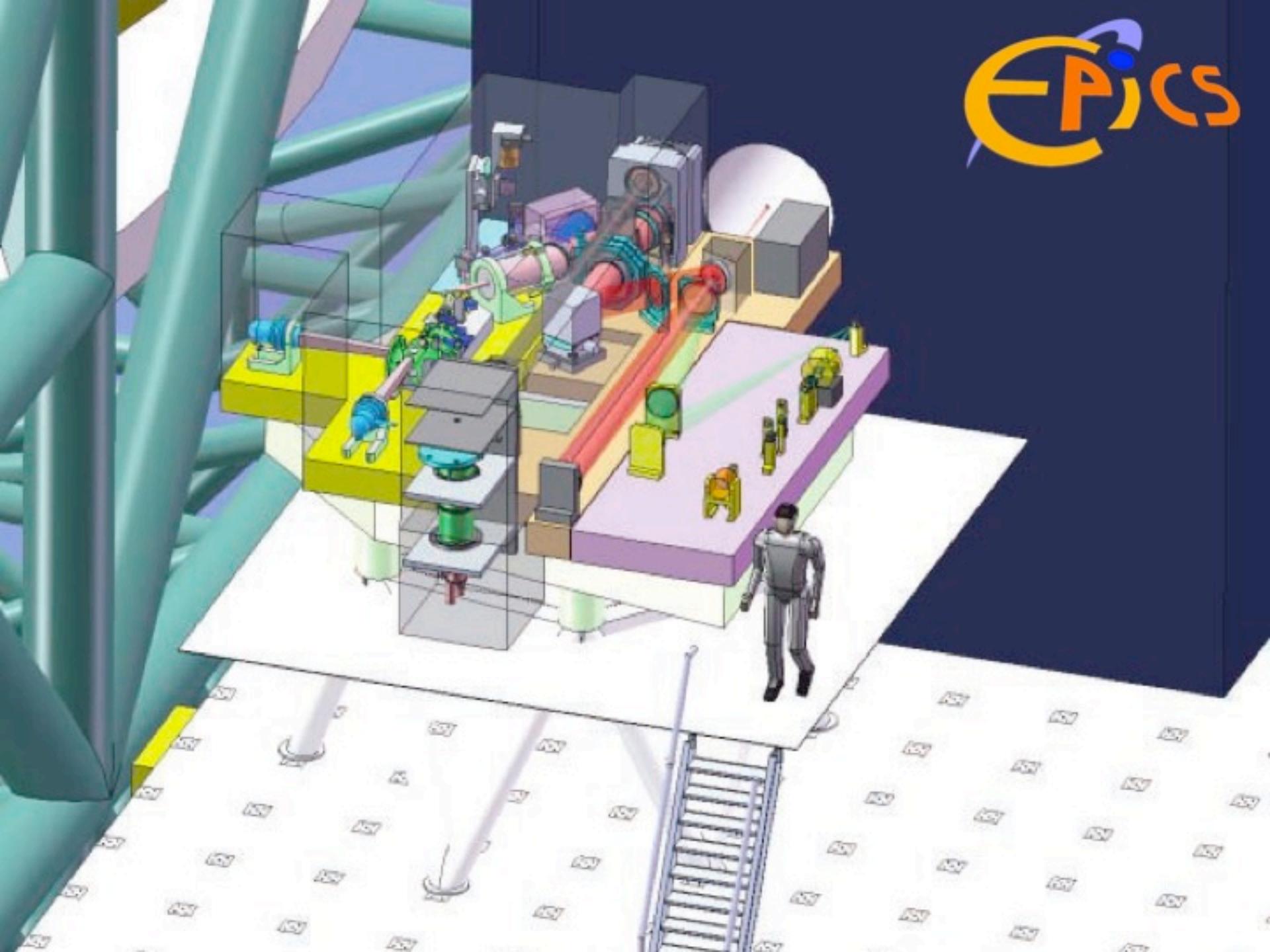


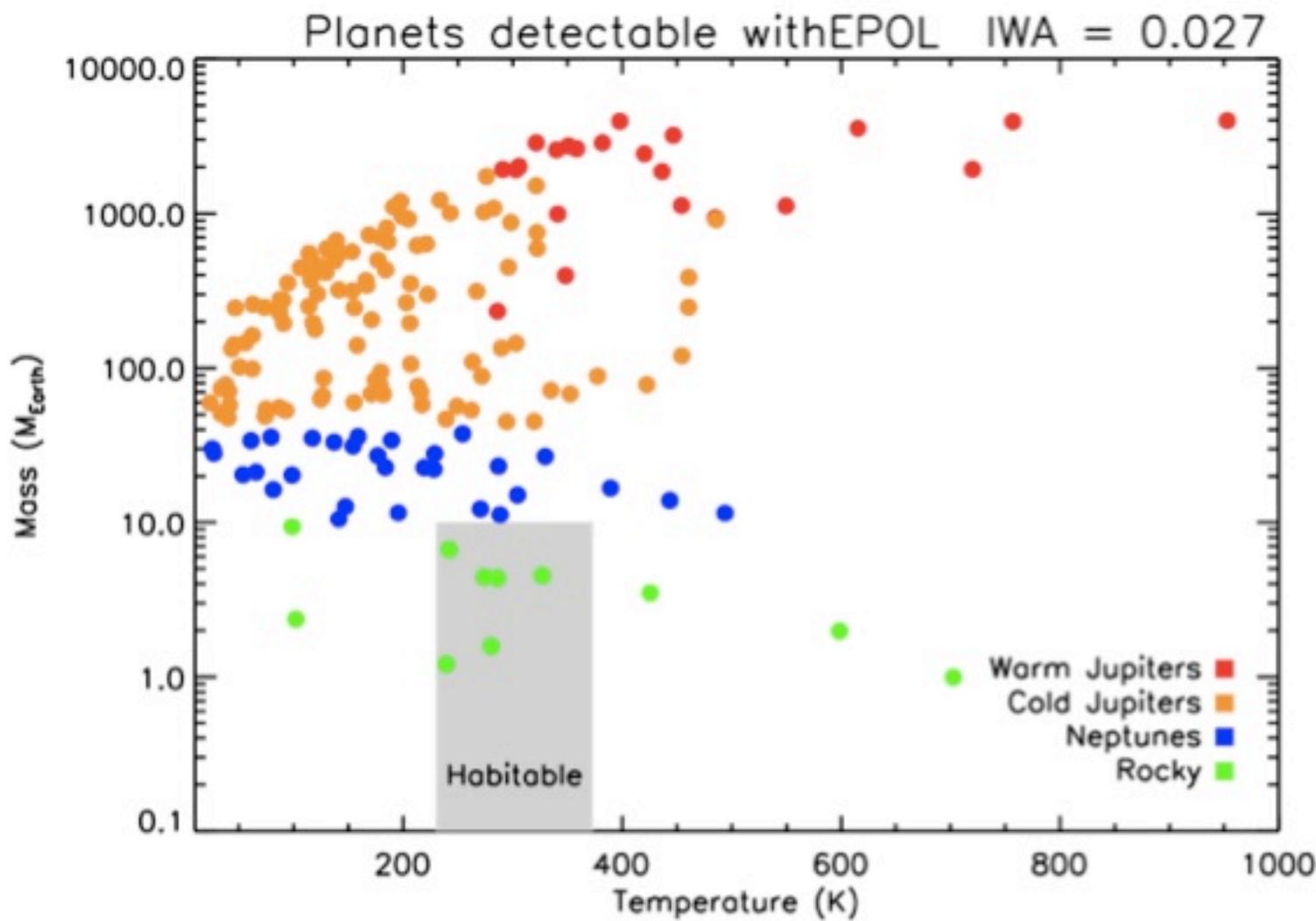
Sphere - ZIMPOL Concept



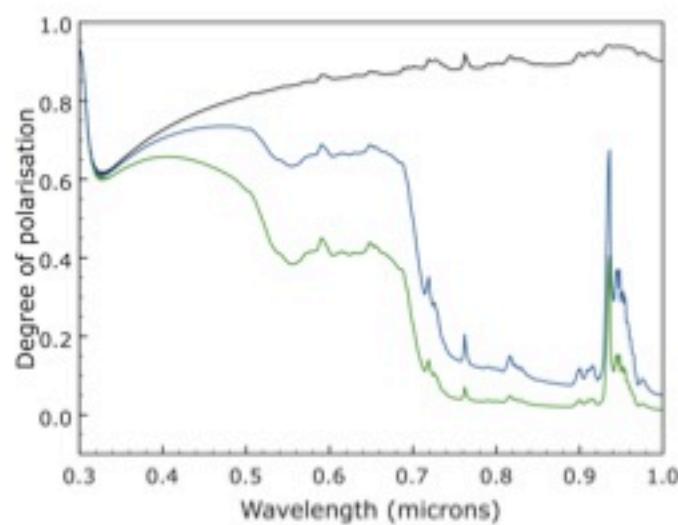
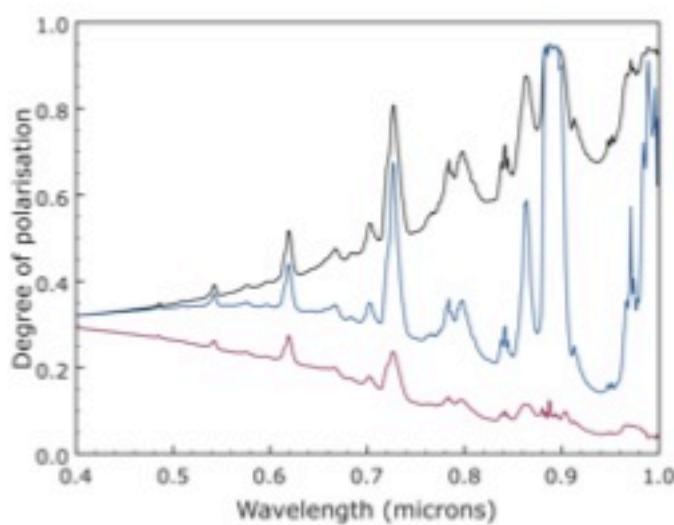
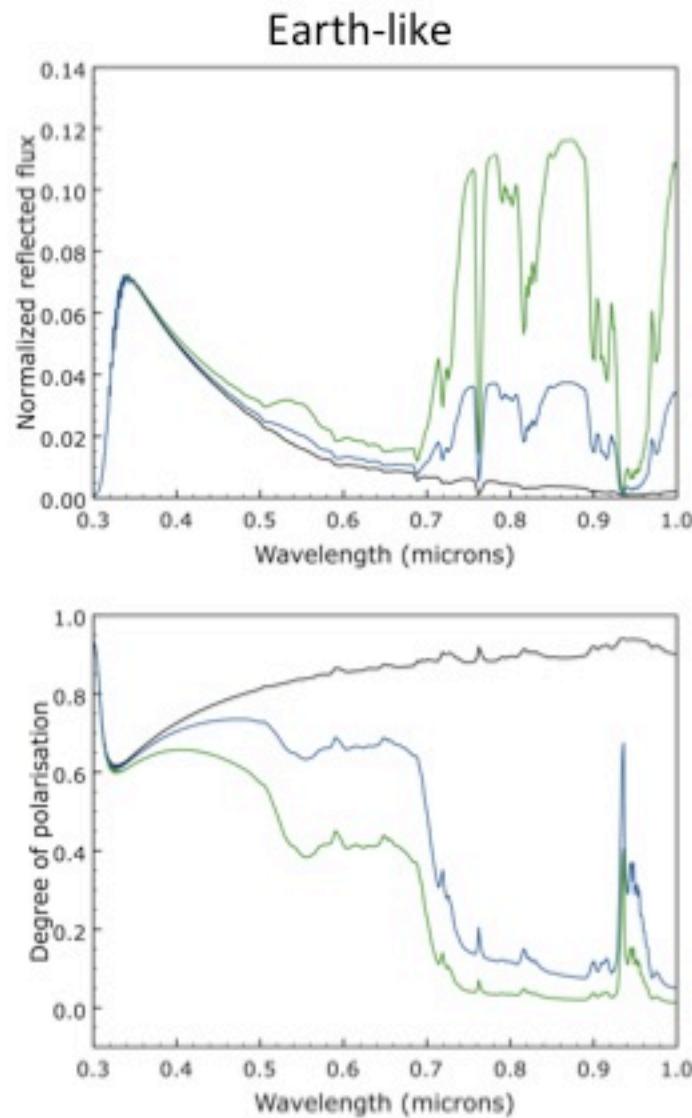
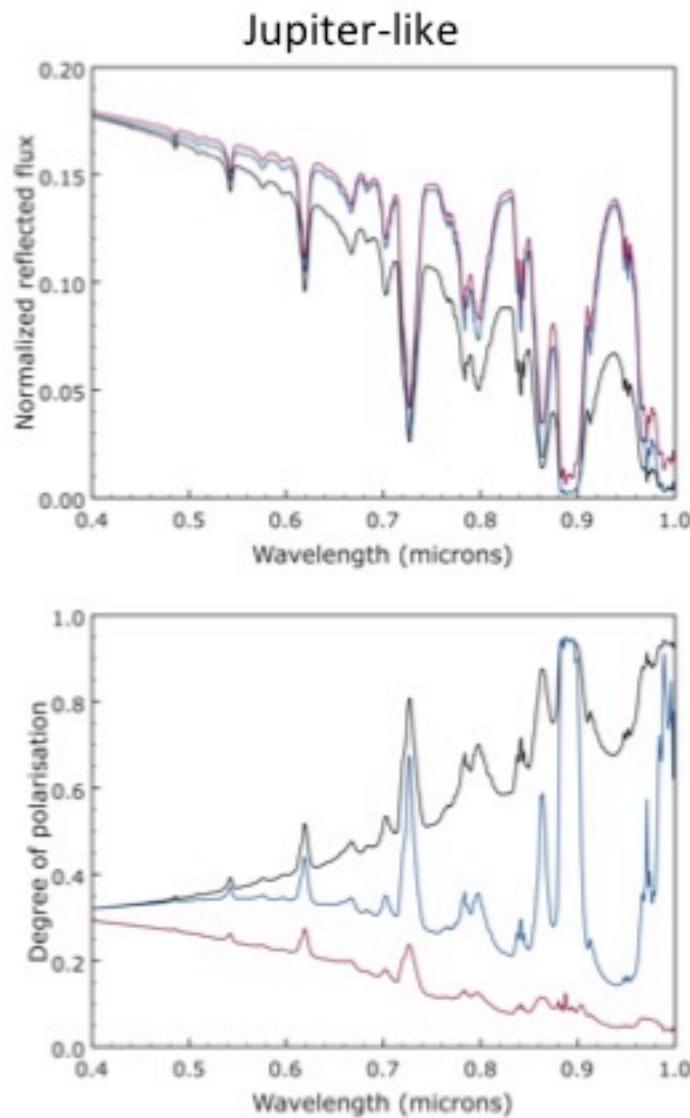
courtesy: Ramon Navarro



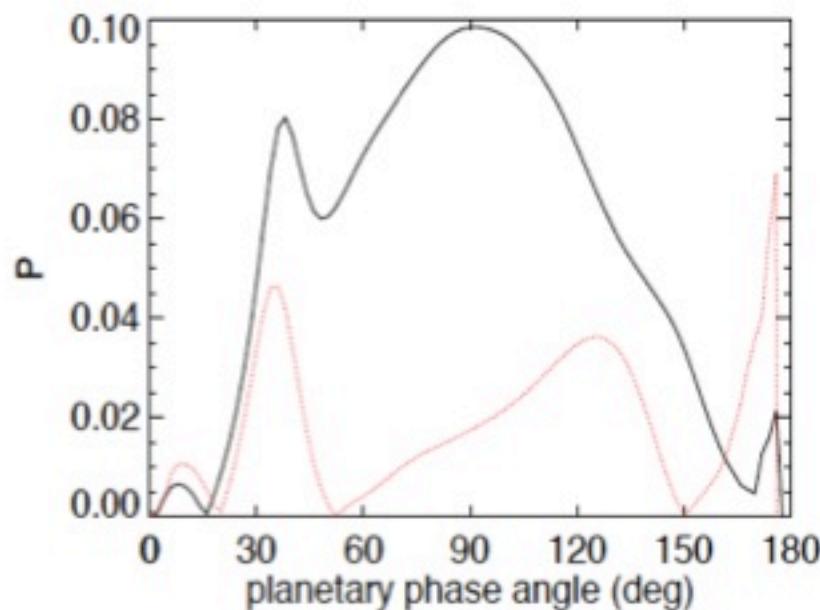
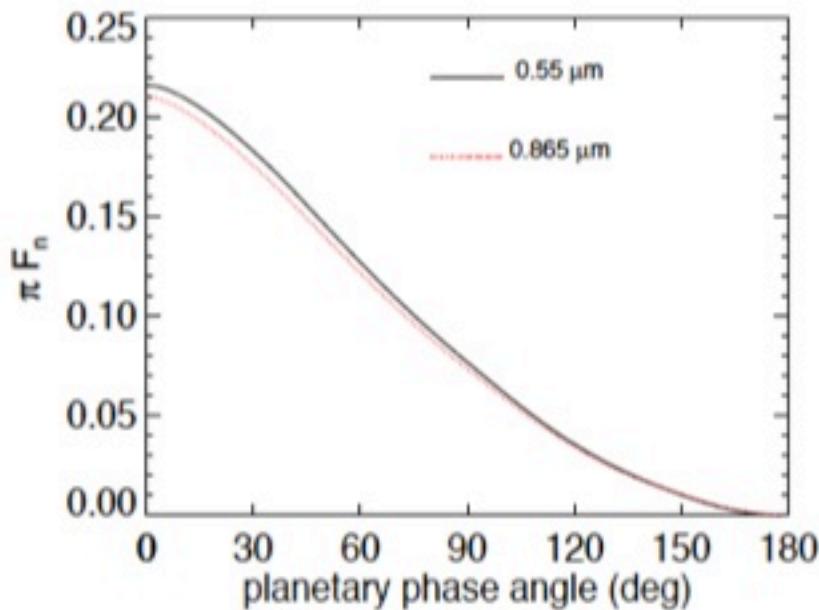
The EPICS logo is located in the top right corner. It consists of the word "EPICS" in a stylized font where the letters are interconnected by a blue curved line. The letters are colored orange, yellow, and blue.



exoplanet characterization



exoplanet characterization



Looking for
extraterrestrial life...

through polarized
glasses!

